



The AMS-02 Experiment Status

Outline

- > News from the past: AMS-01 Ions*
- > The major change happened in July*
- > New testing in August*
- > Present at KSC*
- > (Not-so) Far future*



Istituto Nazionale
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Trieste, 8/9/2010, SciNeGHE 2010

The AMS Physics Goals

CR Hadronic Component

- > Secondary-to-Primary Ratios: CR Propagation Models*
- > Long Period Observations: Solar Modulation*

Indirect Dark Matter Research

- > Anti-particle Spectrum (anti-p, e^+)*
- > γ -ray Flux*

Direct Search for Anti-Nuclei

... and something unexpected

An AMS Brief Description

AMS-01 (1998)

10 Days Data-Taking (Shuttle Mission)

Permanent Magnet ($B = 0.13 \text{ T}$)

Si Tracker ($A = 3 \text{ m}^2, L = 1.2 \text{ m}$) $\rightarrow p, Z$

Time-of-Flight + Veto \rightarrow trigger, β, Z

Aerogel Counters $\rightarrow e/p$ separation

AMS-02 (2011 - ...)

10 Years Data-Taking (ISS Life)

Permanent Magnet ($B = 0.13 \text{ T}$)

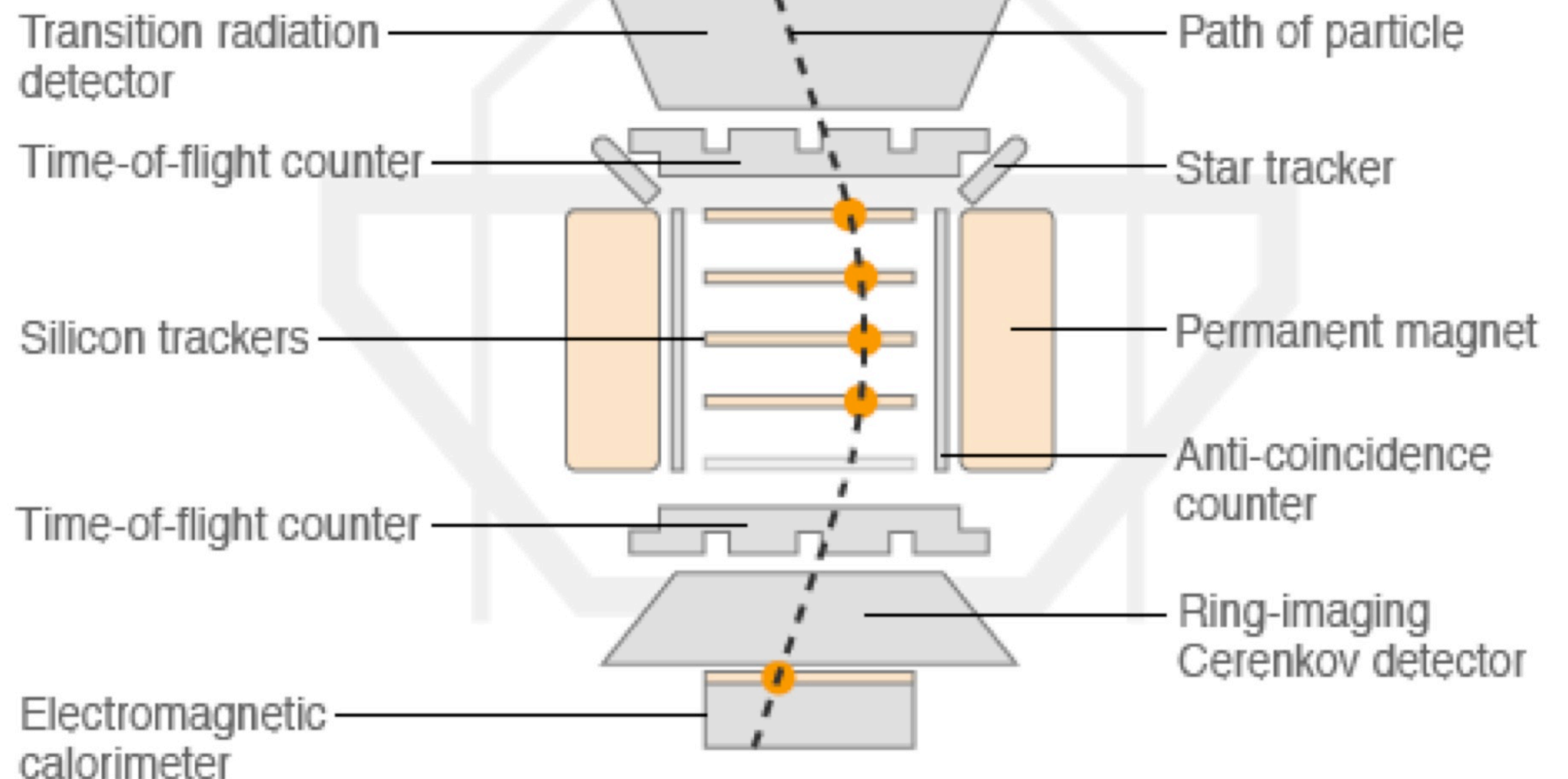
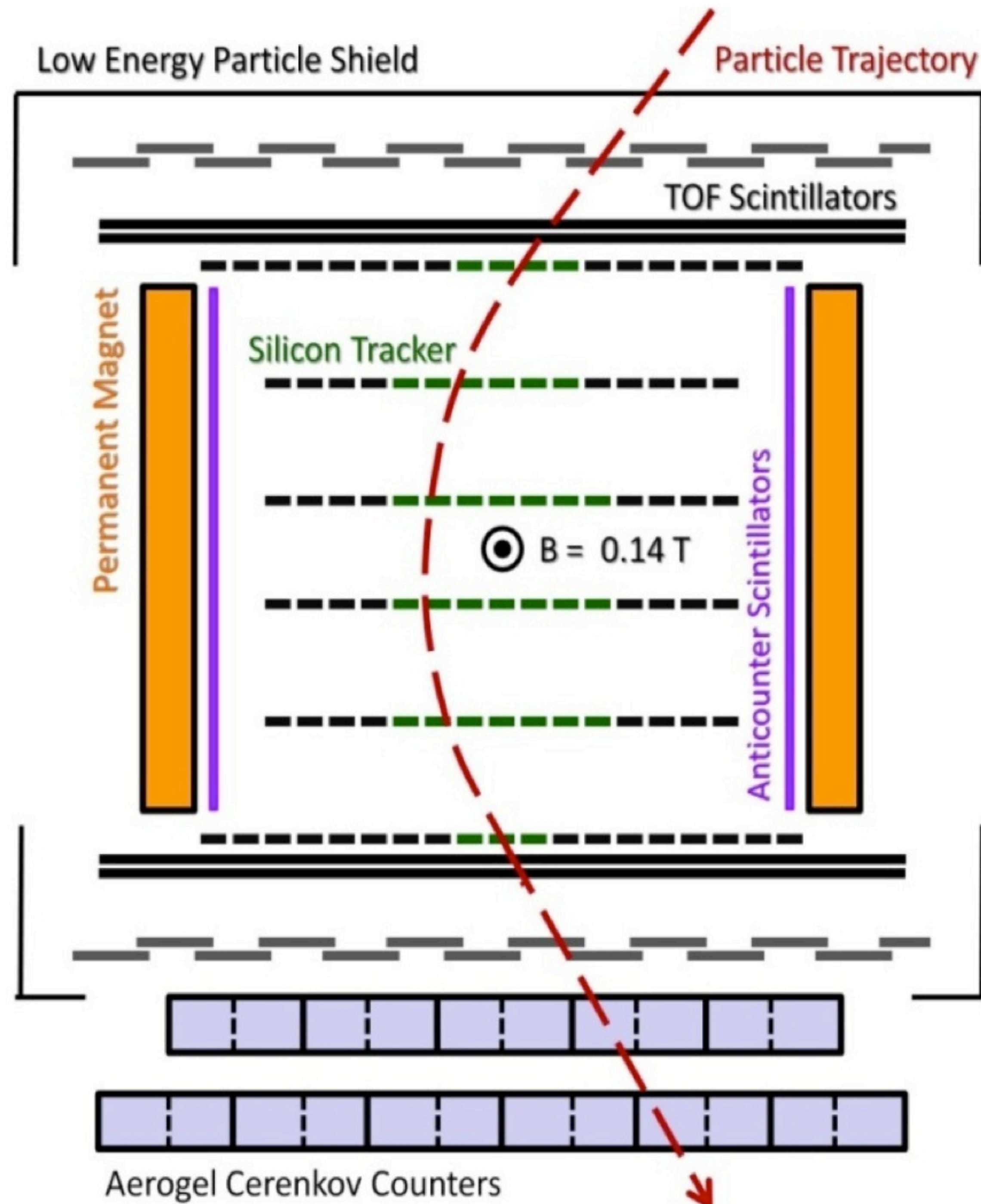
Si Tracker ($A = 6.2 \text{ m}^2, L = 3 \text{ m}$) $\rightarrow p, Z$

Time-of-Flight + Veto \rightarrow trigger, β, Z

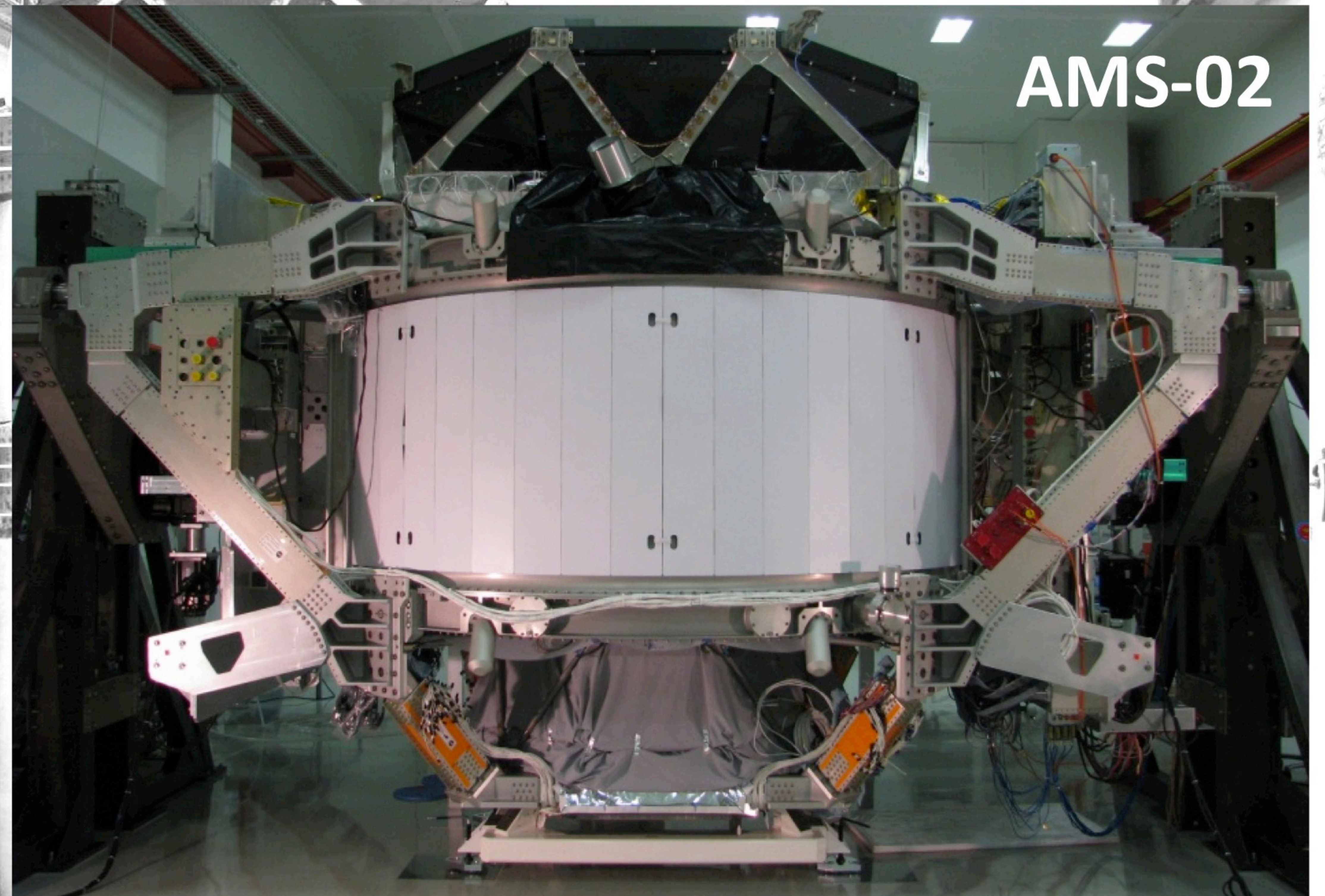
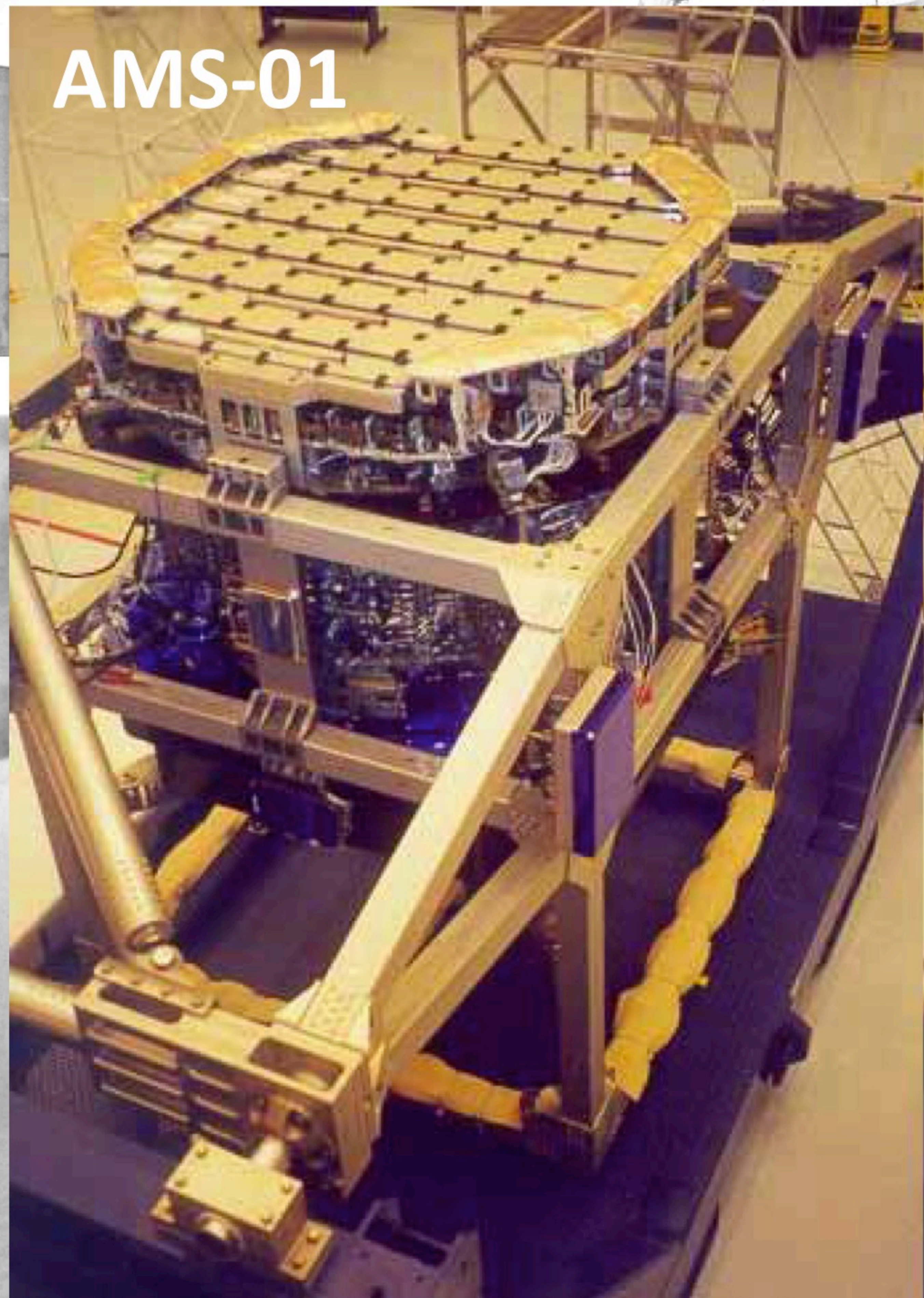
Ring Imaging Cherenkov $\rightarrow \beta, Z$

Transition Radiation Detector $\rightarrow e/p$ sep.

Electromagnetic Calorimeter $\rightarrow e/p$ sep., E of e & γ



AMS-01 and AMS-02

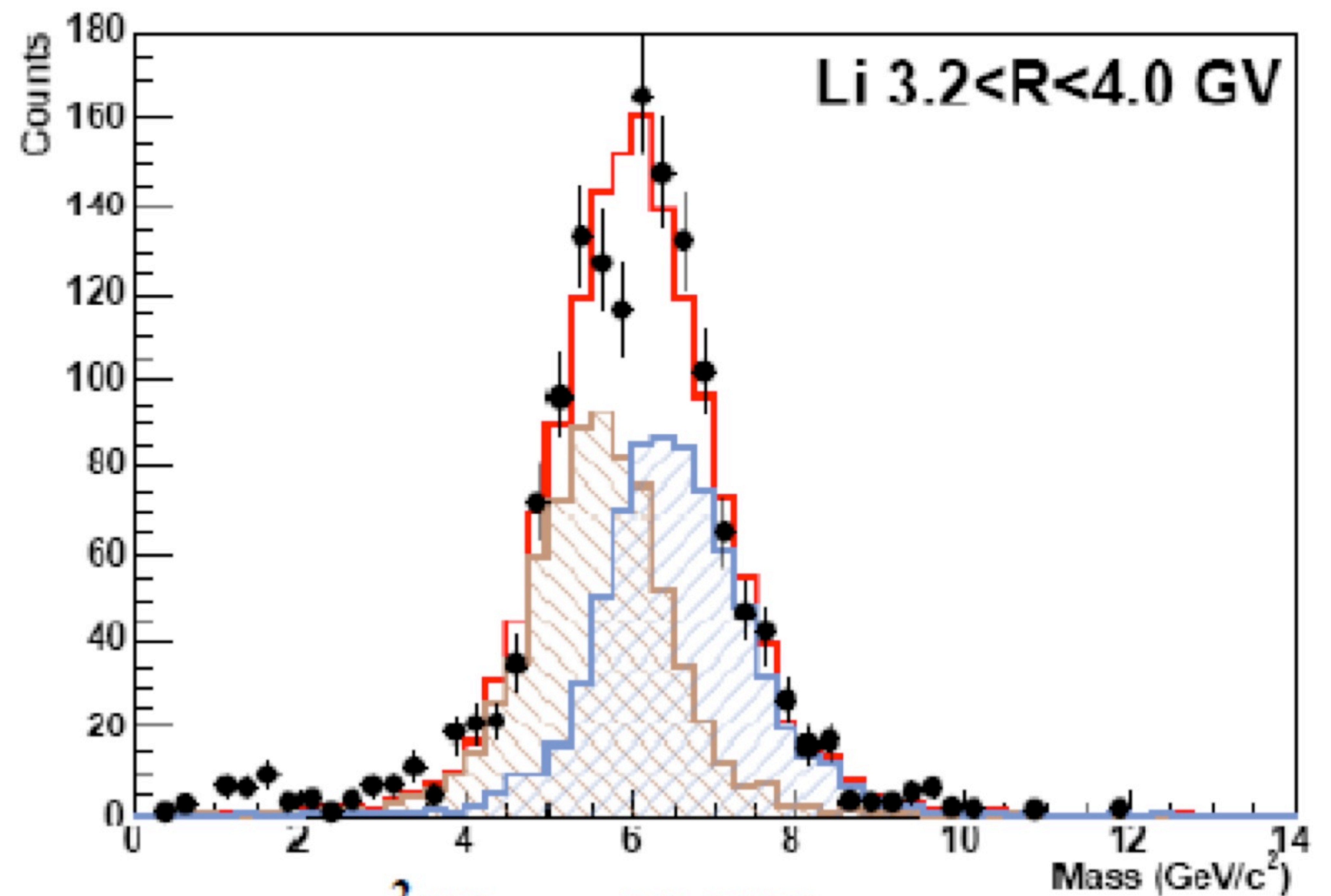
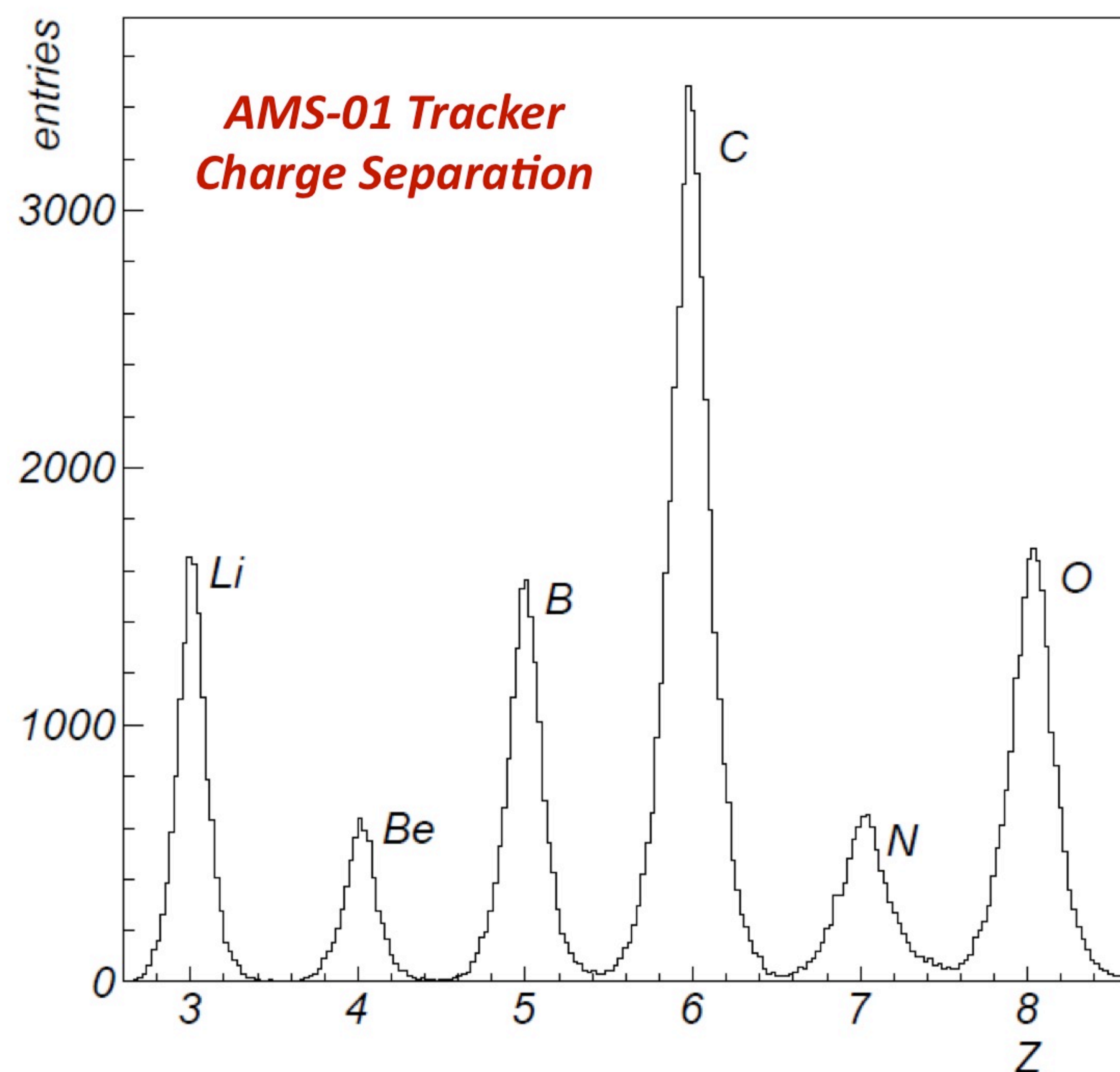


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A New Result: The AMS-01 CRNs Analysis

Most of our knowledge of Cosmic Rays comes from Cosmic Rays Nuclei (CRN).
Propagation models depend from the secondary-to-primary ratios such as Li-Be-B/C.



Chemical Composition

Si Tracker Z Res. = 0.10 (Li) to 0.14 (C) c. u.

Identification Eff. \approx 99%

Contamination on B/C = $2 \div 9\%$

He-Li separation from TB data (ToF + Tracker)

Isotopic Composition

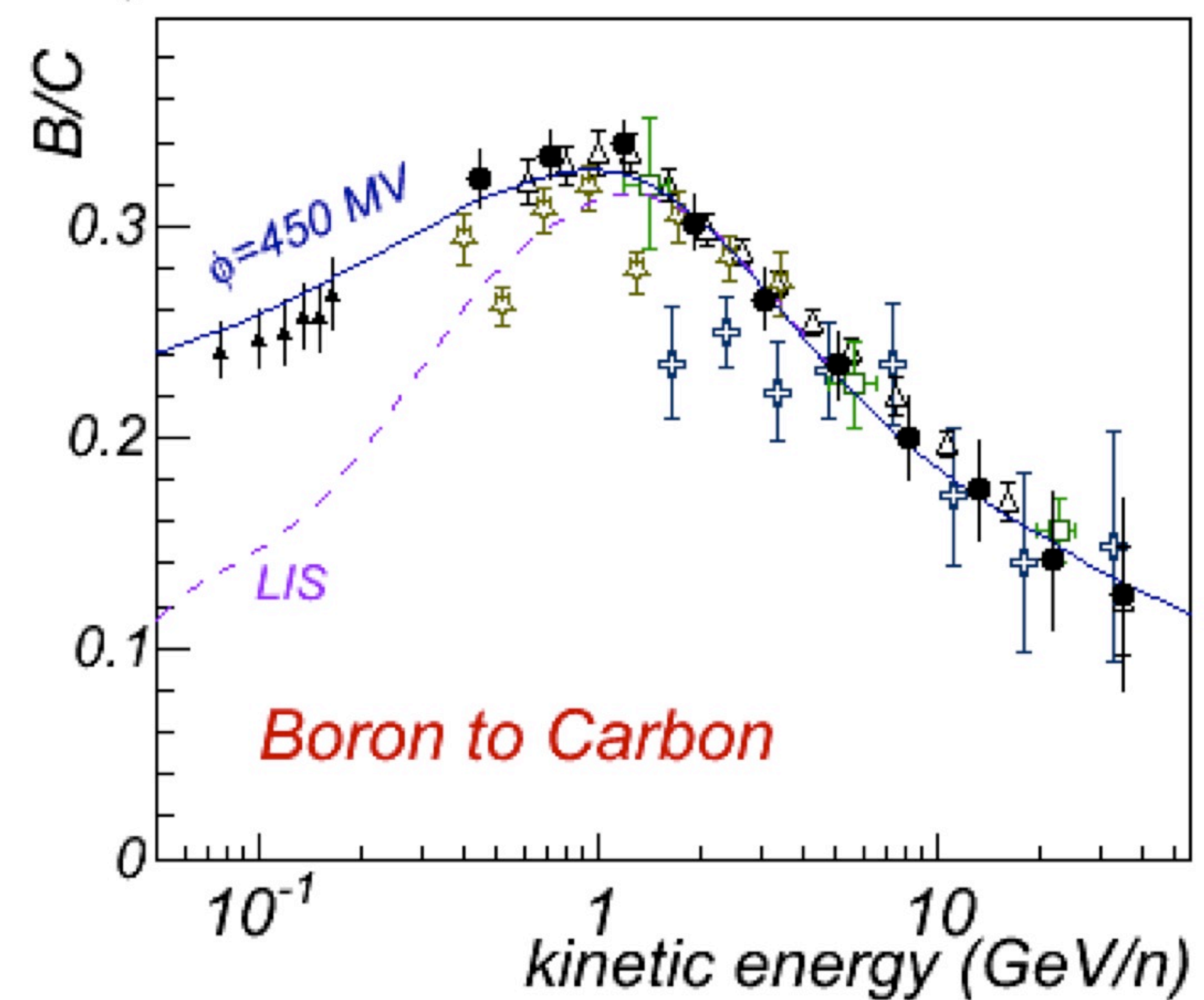
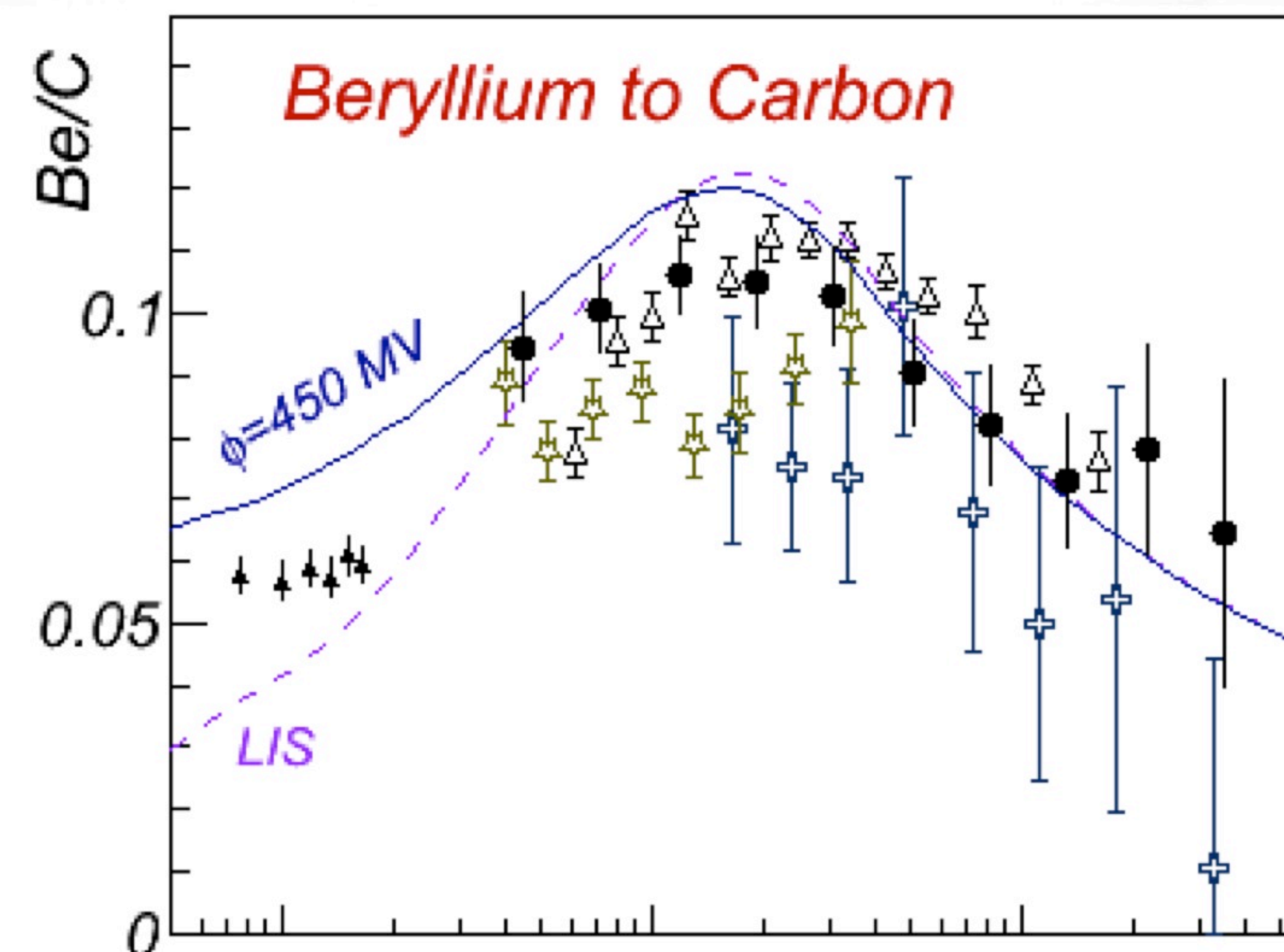
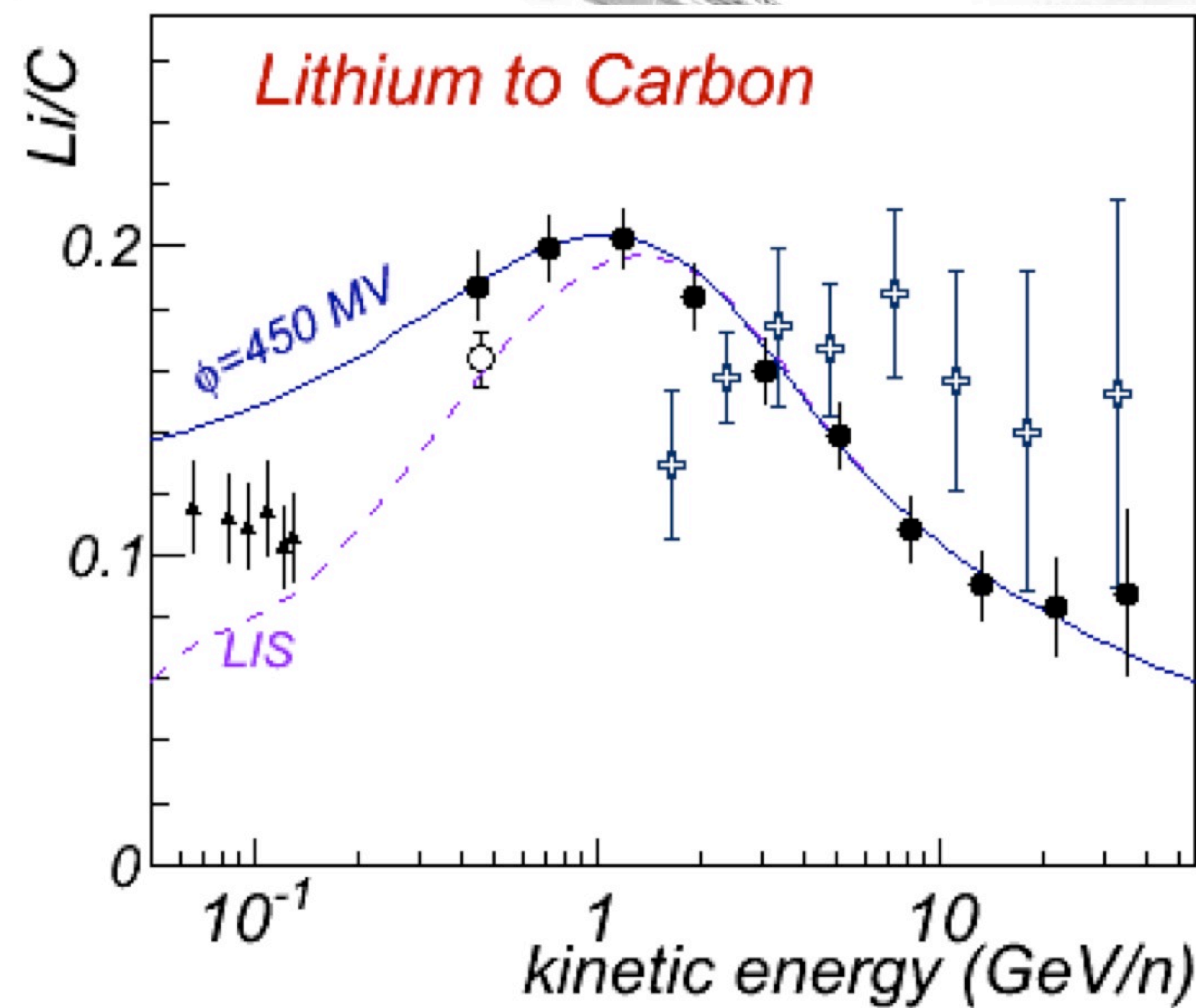
Mass from Tracker + ToF (up to R < 6.3 GV)

Composition Fit Technique (use of MC)

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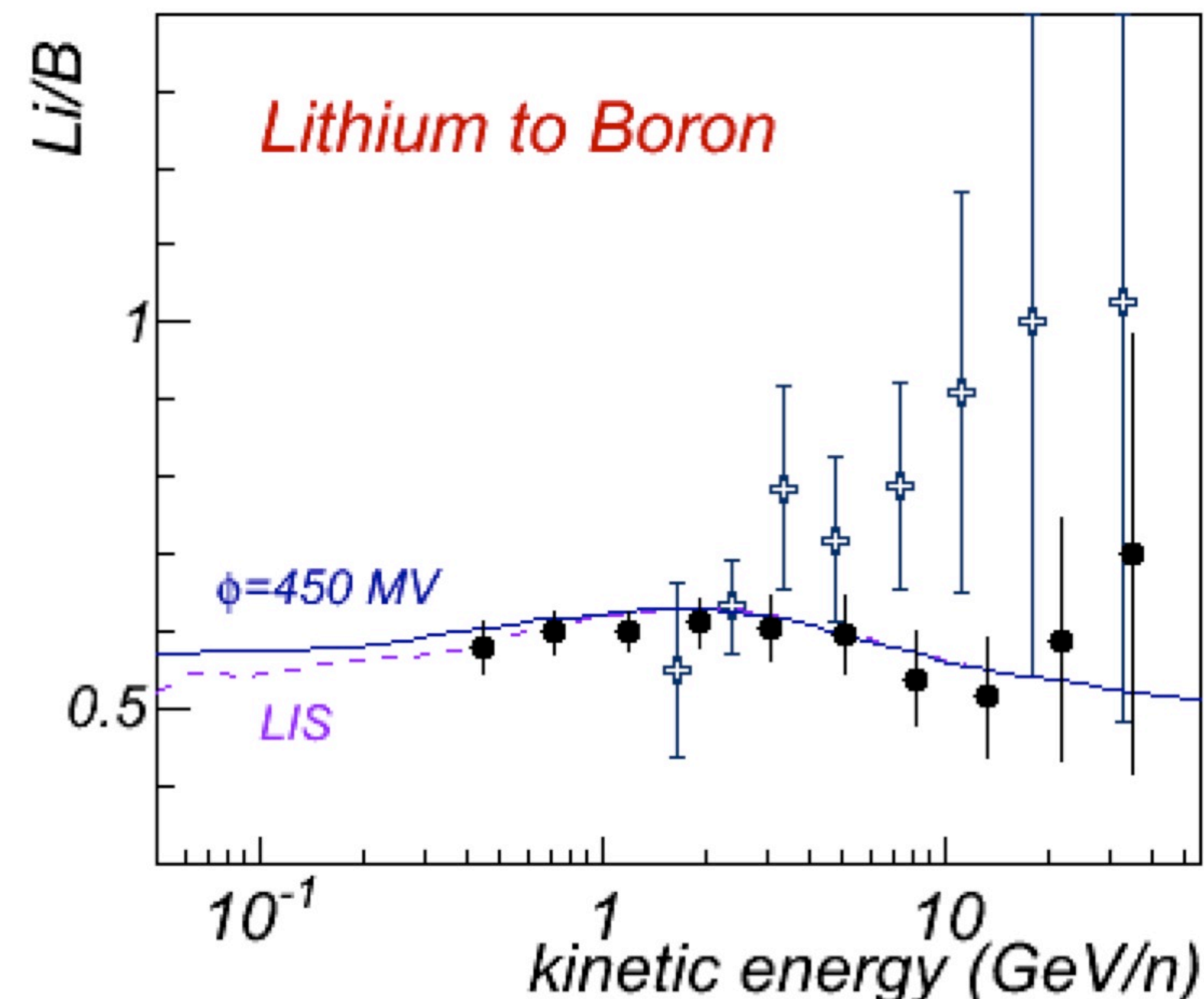
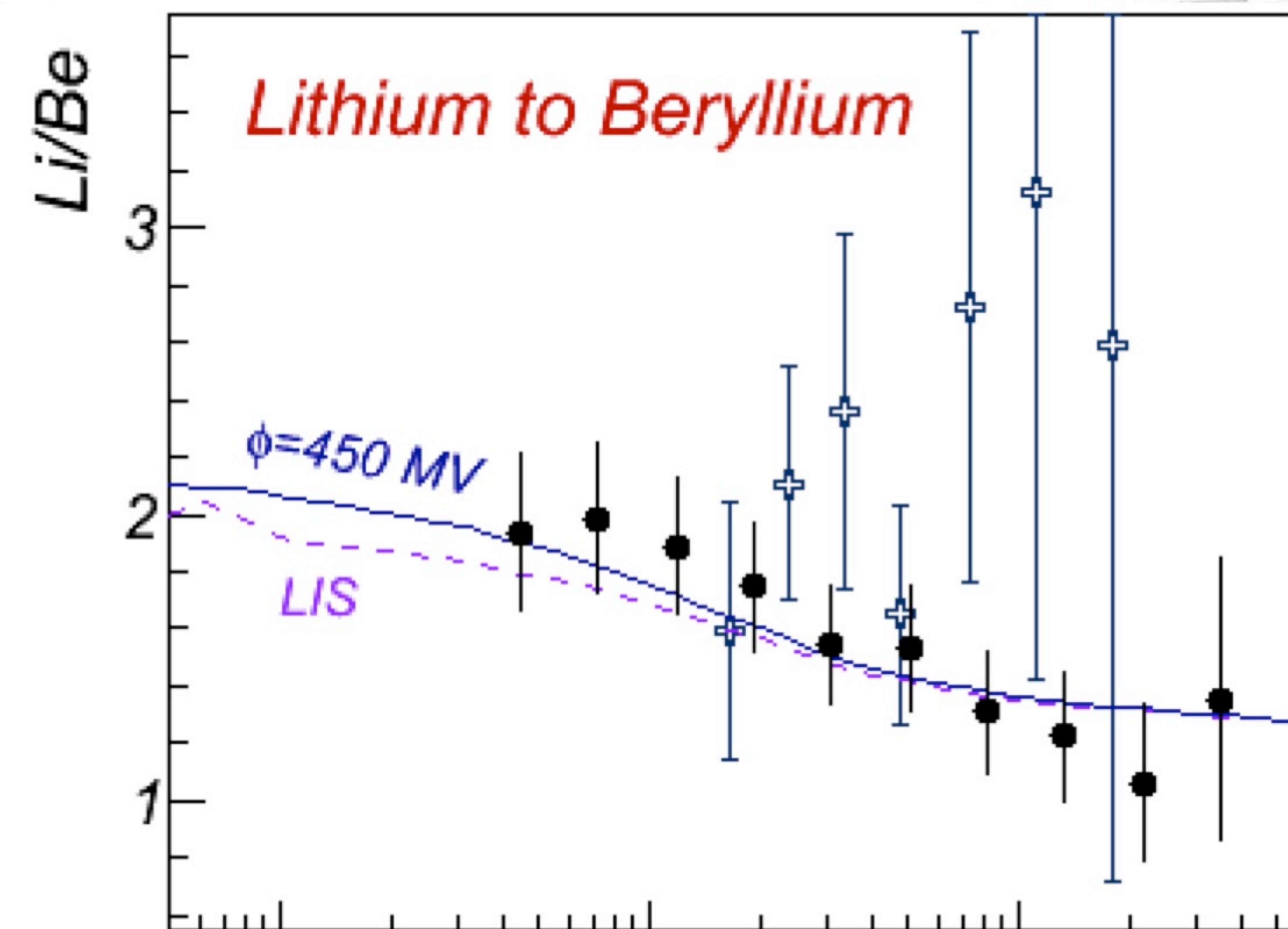
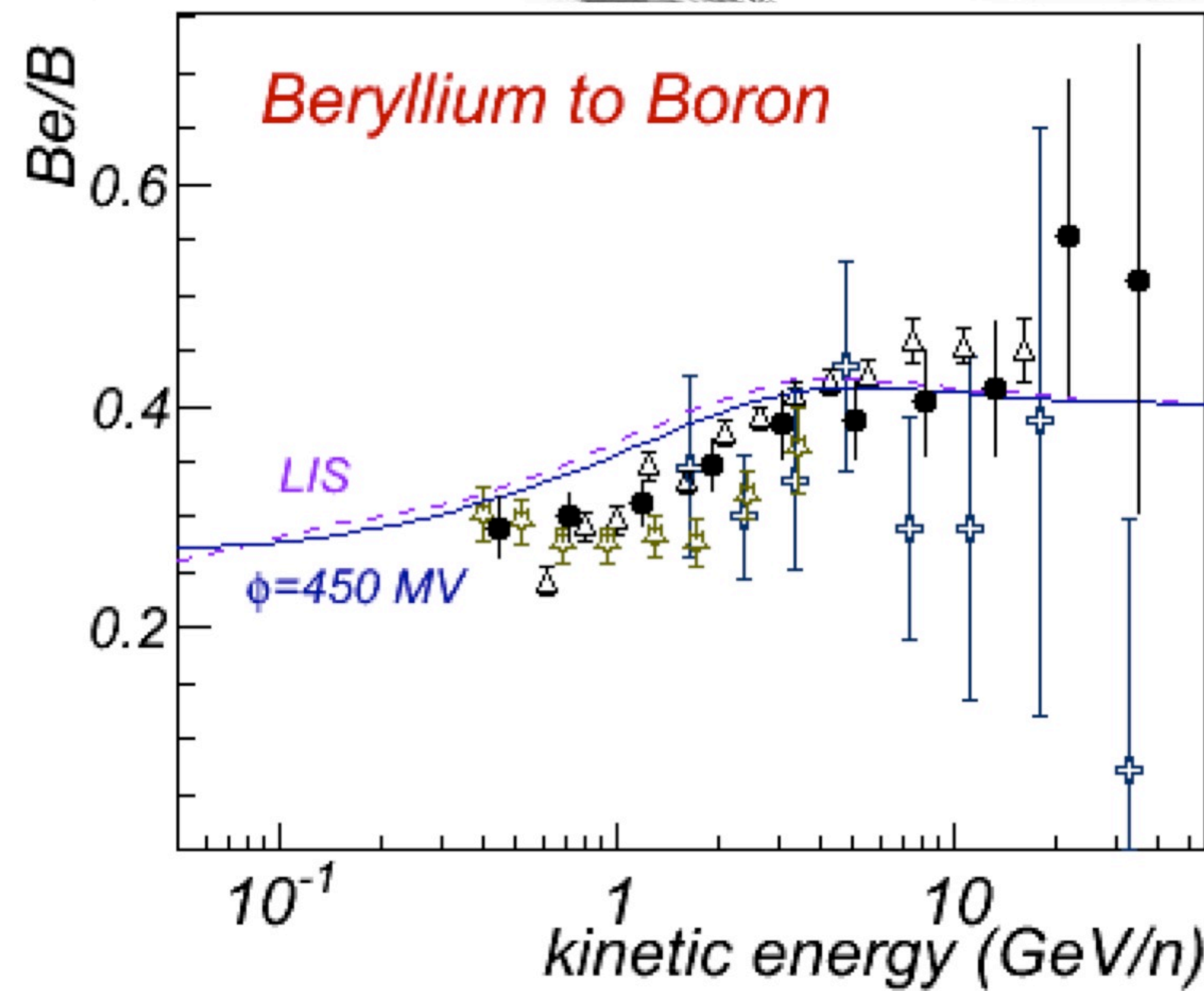
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The AMS-01 CRNs Analysis: *Secondary-to-Primary*

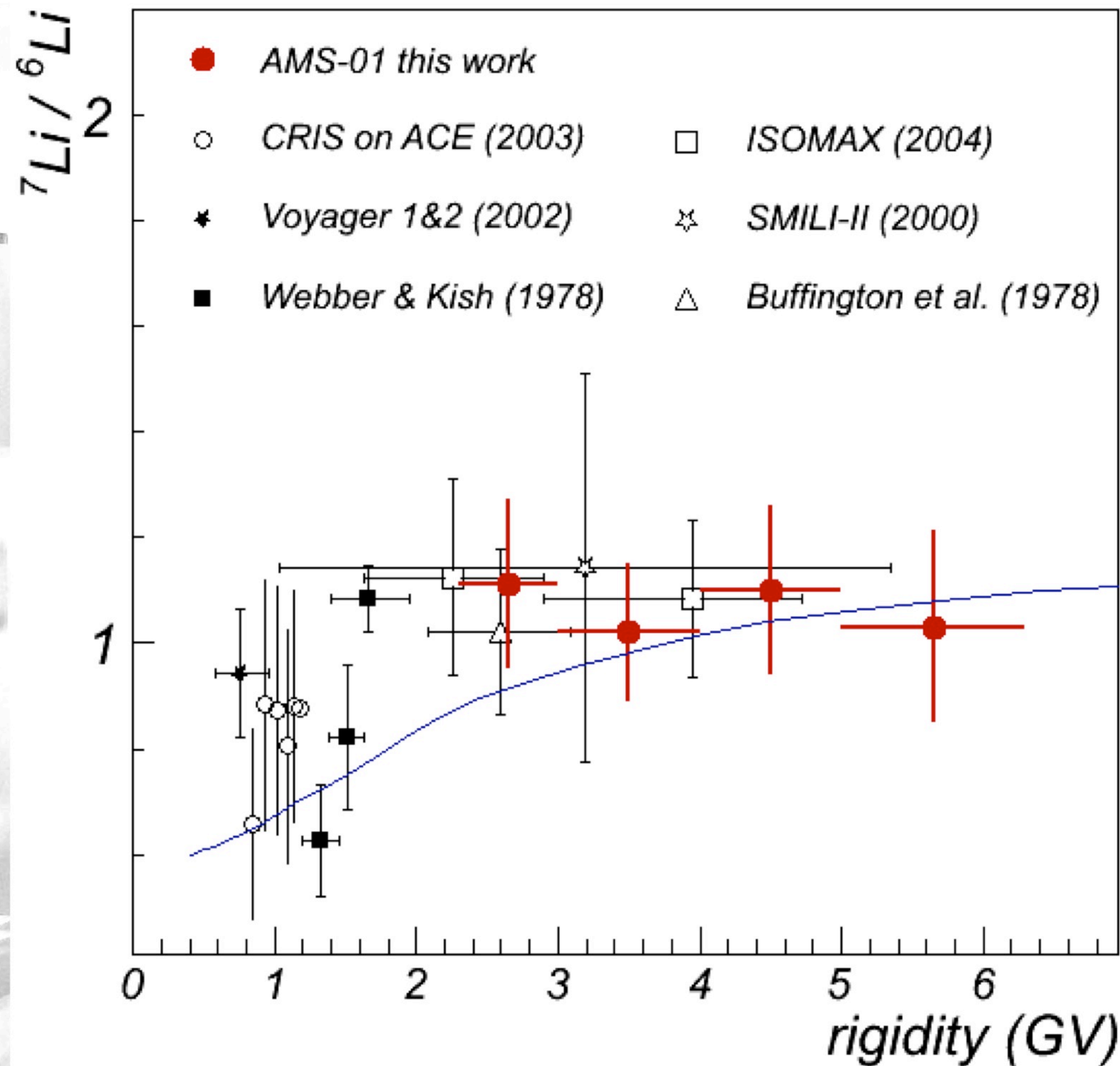


- AMS-01 this analysis
- Webber et al. 1972
- ⊕ Orth et al. 1978
- △ HEAO3-C2 1990
- CREAM 2008
- ☆ Lezniak & Webber 1978
- ▲ CRIS/ACE 2003

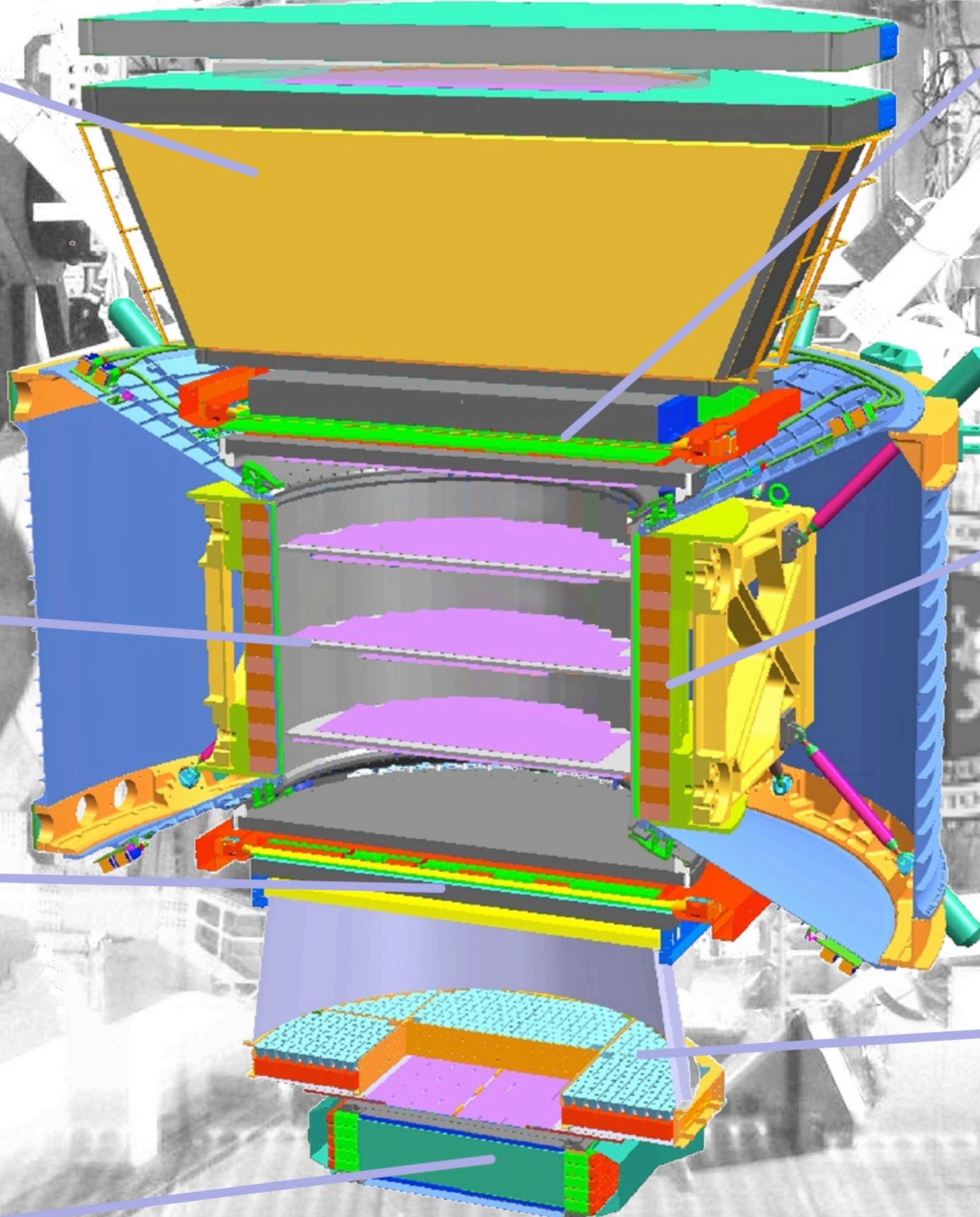
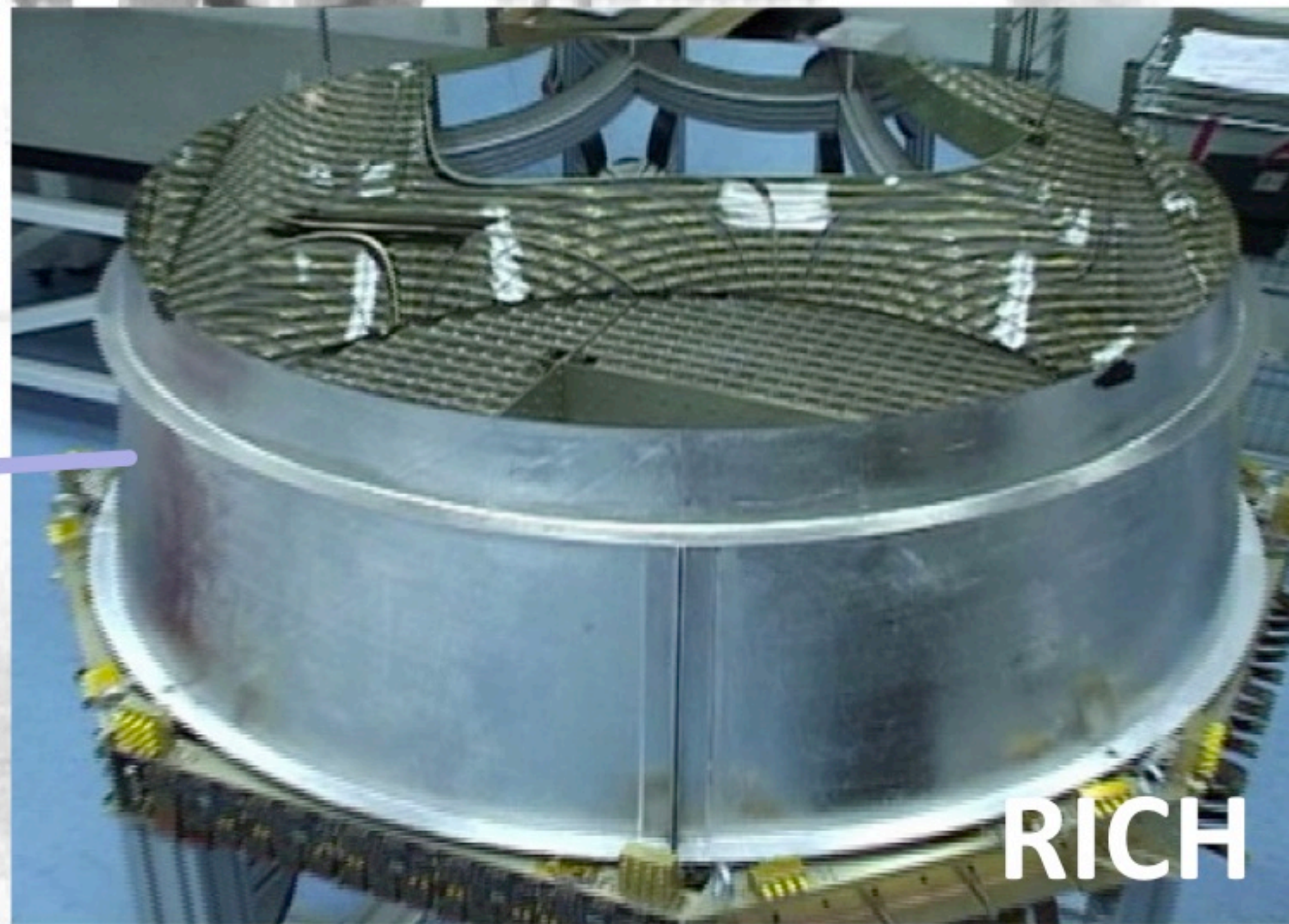
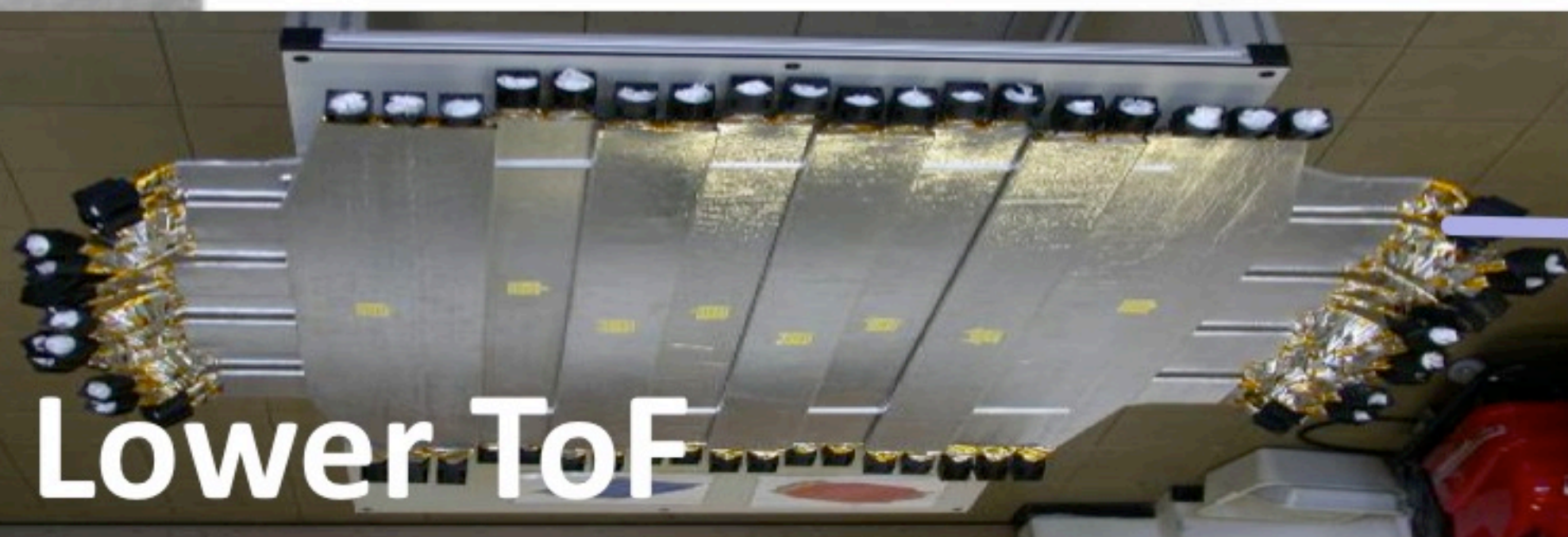
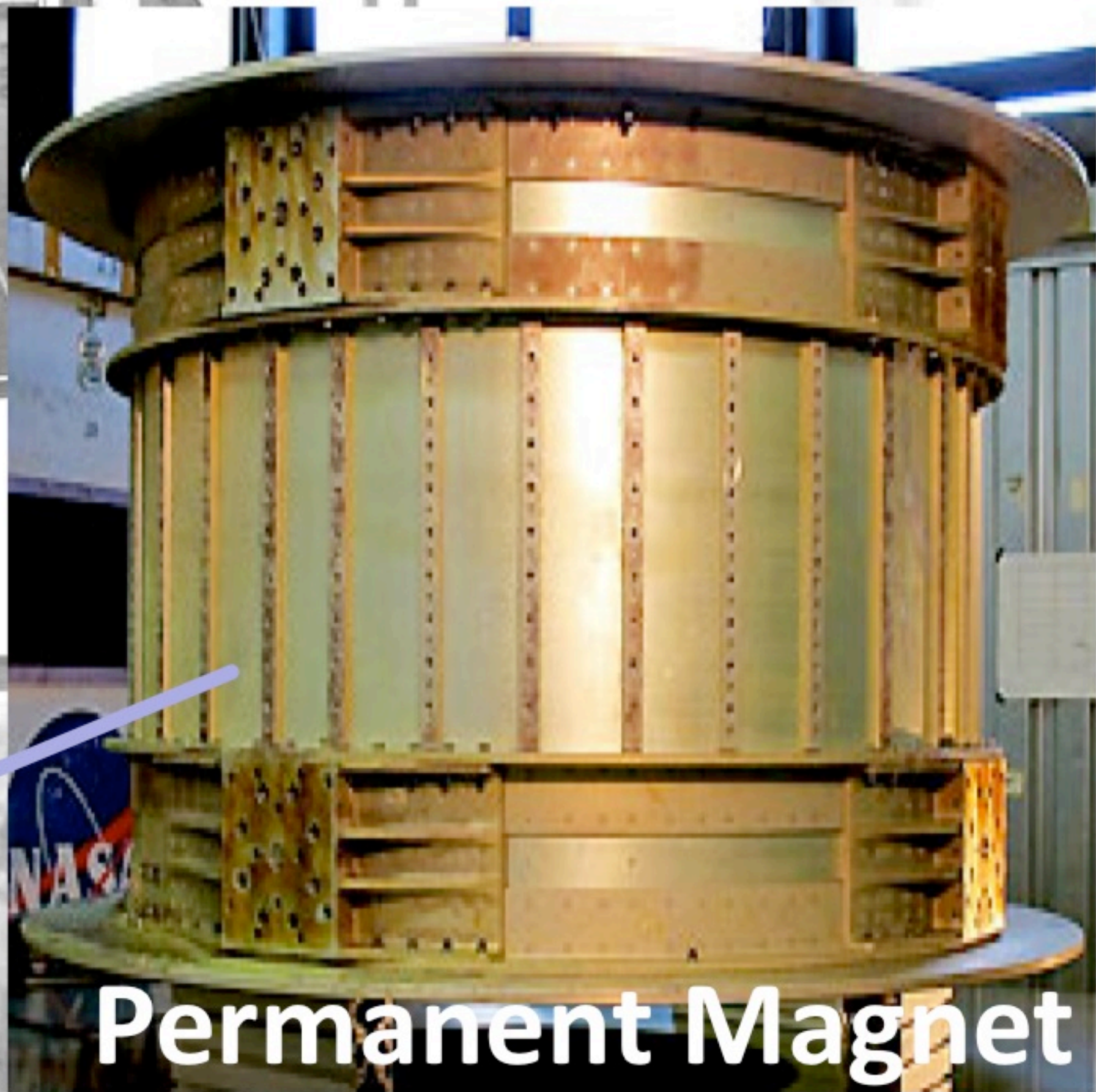
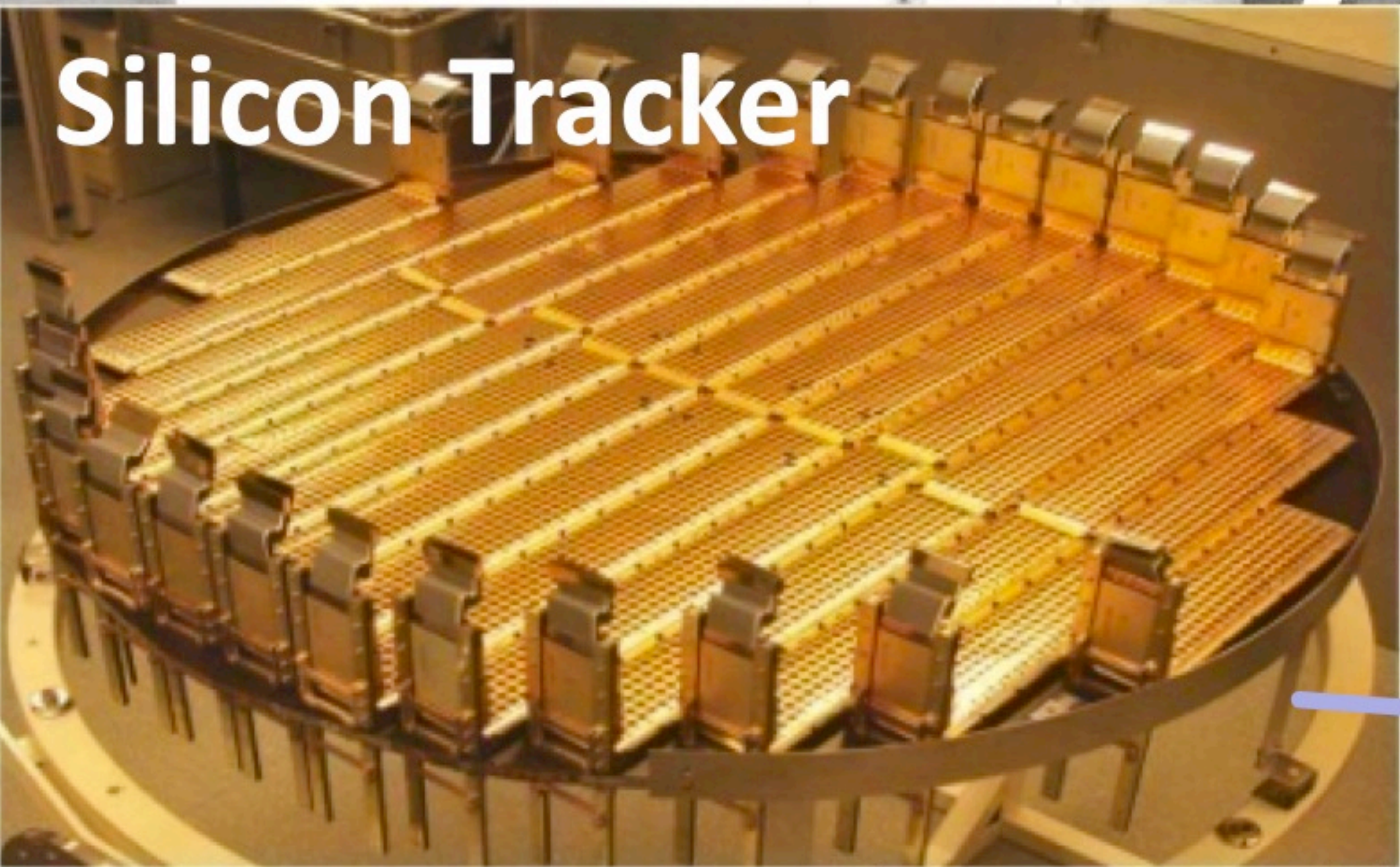
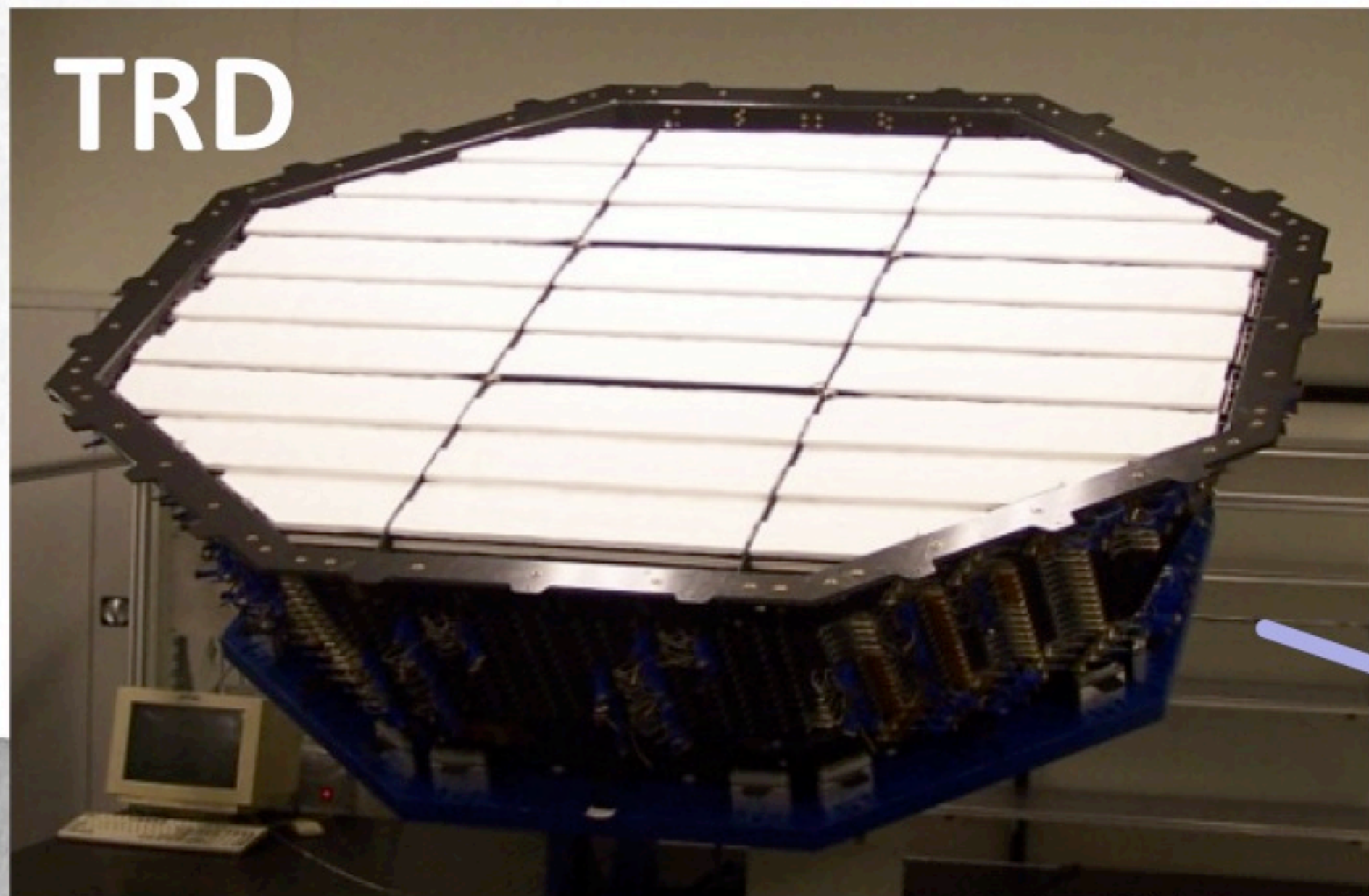
The AMS-01 CRNs Analysis: *Secondary-to-Secondary*



The AMS-01 CRNs Analysis: *Isotopic Ratios*



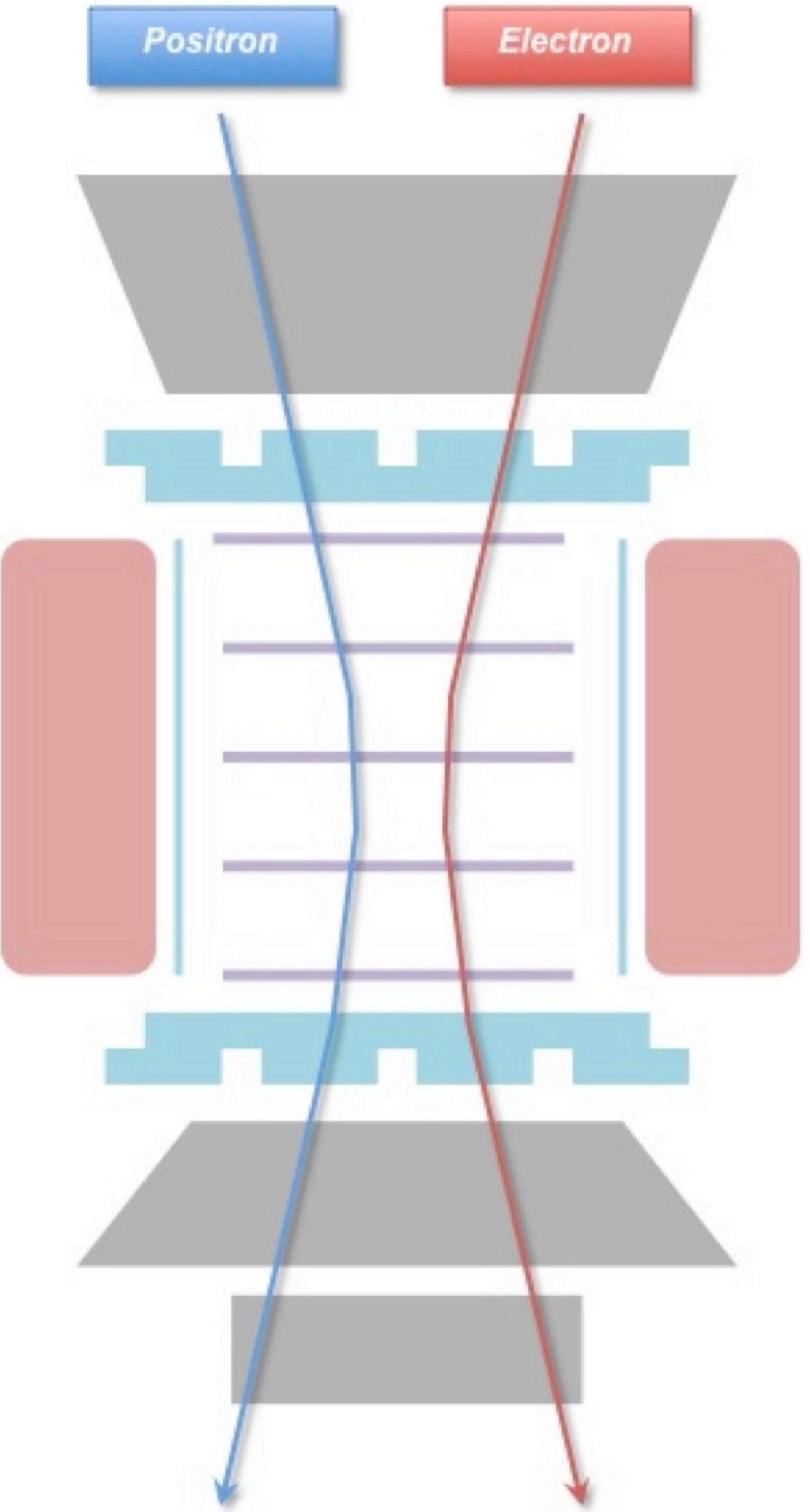
AMS-02 Subsystems



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AMS-02 Measurements



	e^-	p	Fe	e^+	\bar{p}	\overline{He}	γ
TRD	↓ ↓ ↓	∇	∇	↓ ↓ ↓	∇	∇	∇
TOF	∇	∇ ∇	∇ ∇	∇	∇ ∇	∇ ∇	
Tracker + Magnet	∪	∩	∩	∪	∪	∩	∪
RICH	○	○	●	○	○	●	
ECAL	∧	∇ ∇ ∇	∇ ∇ ∇	∧	∇ ∇ ∇	∇ ∇ ∇	∧
	Cosmic Rays			Dark Matter		Antimatter	γ Sky

AMS-02 Measurements

Particle Sign and Rigidity (\pm , R)

ToF: up-going to down-going distinction (10^9 rej.)

Magnet: bending

Tracker: measures the deflection up to MDR (2 TeV)
 from orientation \rightarrow charge sign
 from curvature radius \rightarrow momentum

Absolute Charge (Z)

ToF: up to 4 dE/dx ($\propto Z^2$) independent meas.

Tracker: up to 9 dE/dx ($\propto Z^2$) independent meas.

RICH: N_γ ($\propto Z^2$) measurement
 Expected complete separation

Velocity

ToF: $\Delta\beta/\beta \approx 3\%$ ($\Delta t \approx 160$ ps, $L = 1.2$ m)

RICH: $\Delta\beta/\beta \approx 0.05$ (0.01) % for p (ions)

e/p Separation

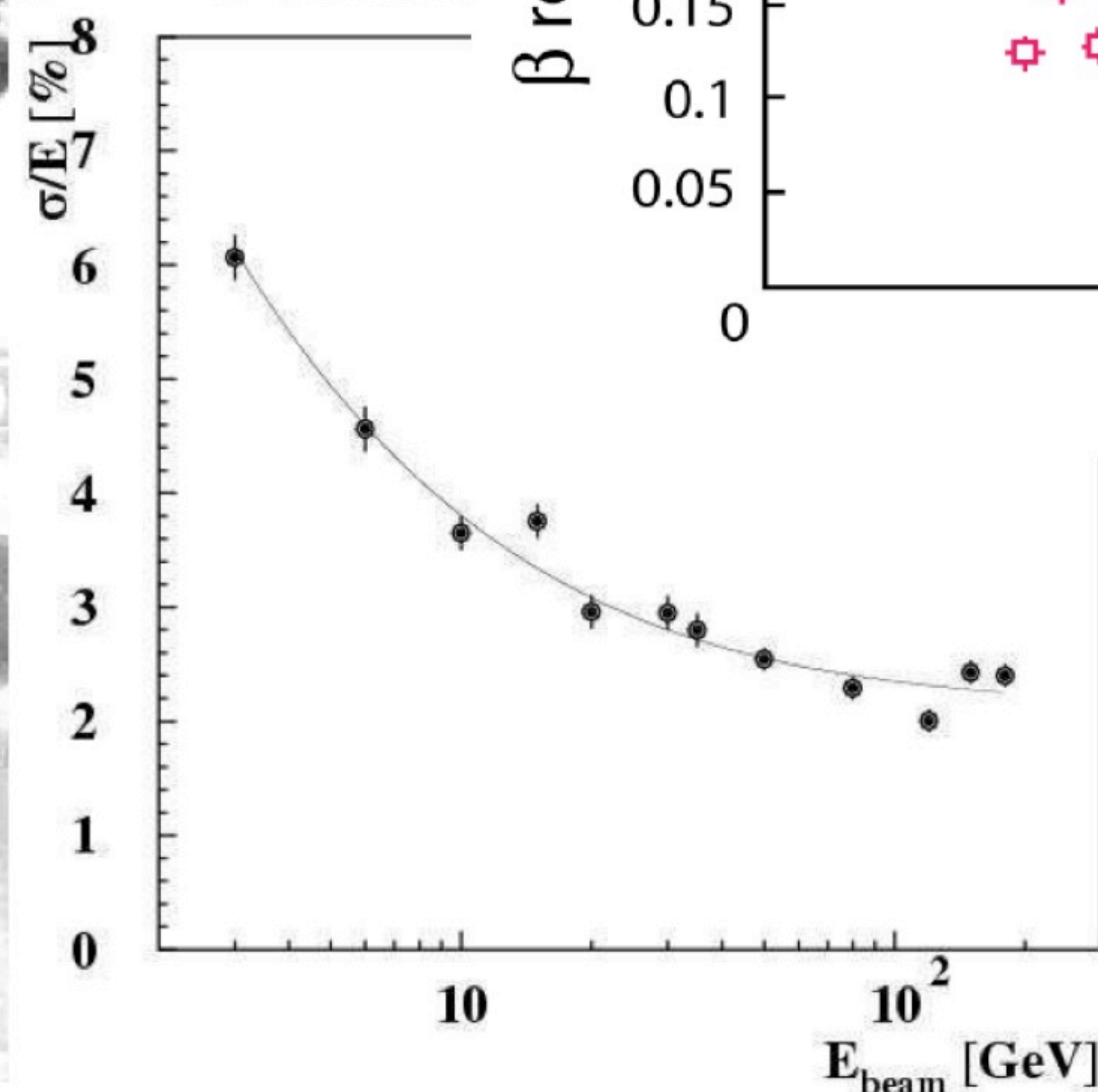
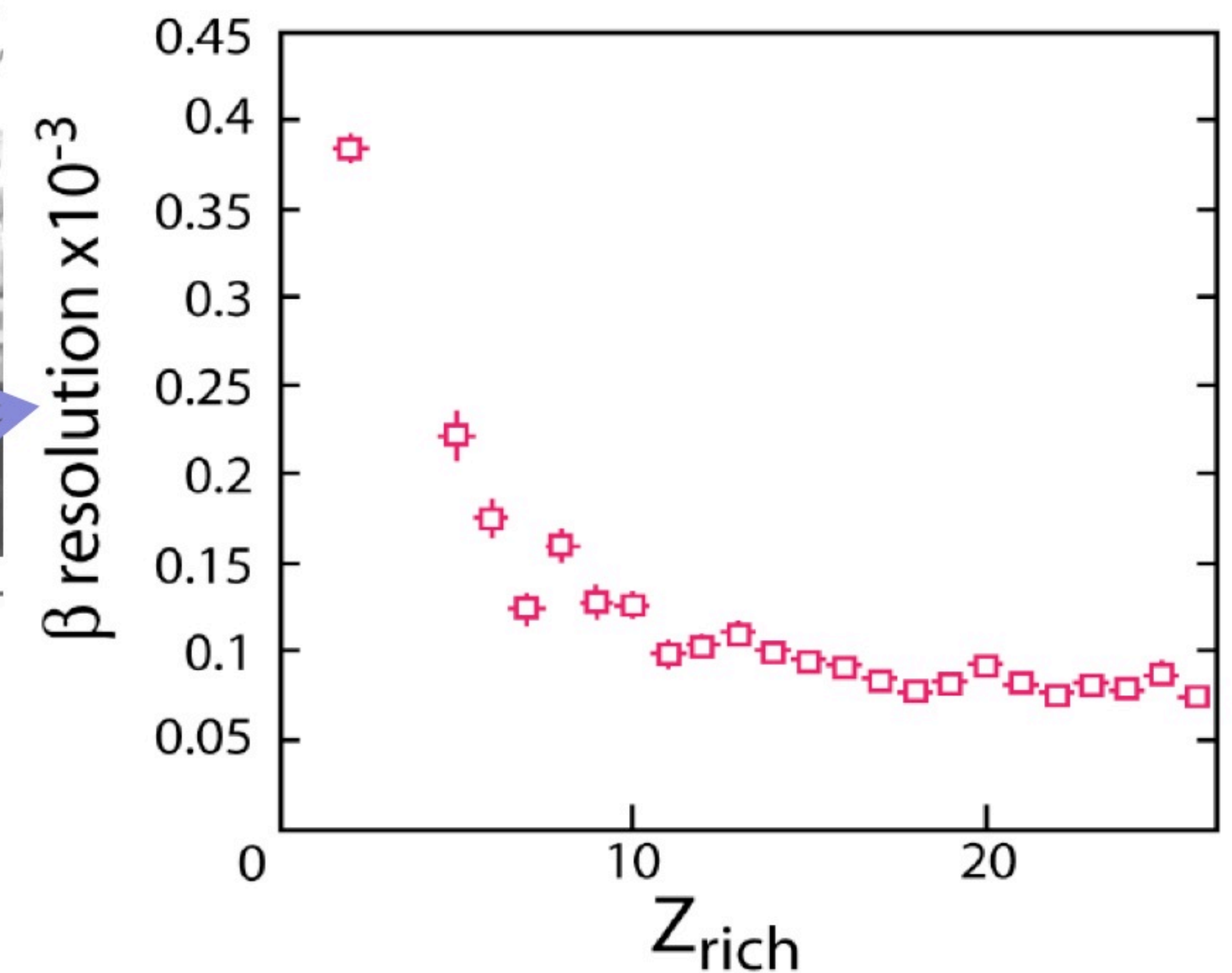
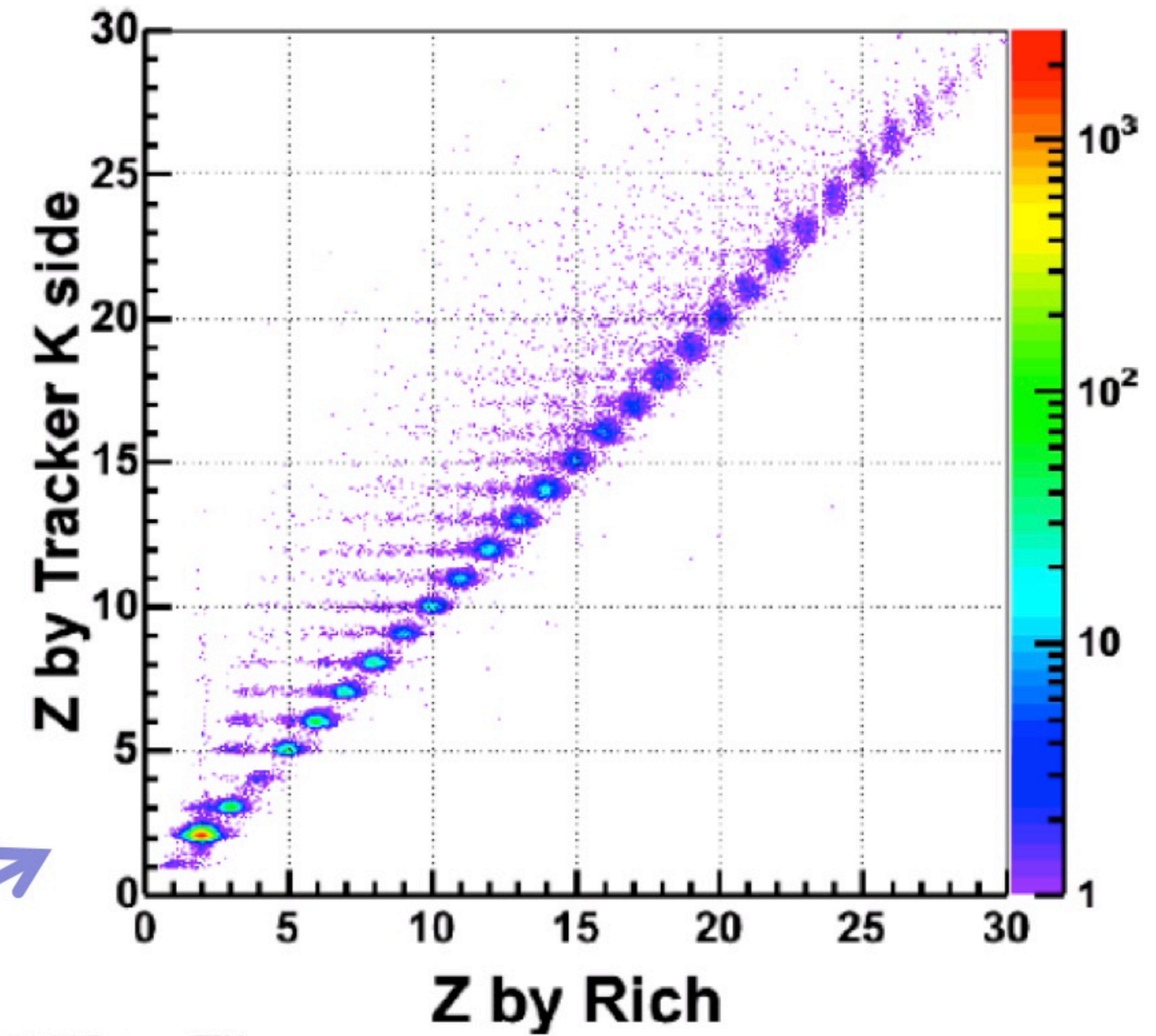
TRD: 10^3 p rejection. up to $E \approx 300$ GeV

ECAL: 10^4 p rejection

Energy and γ

ECAL: $\Delta E/E \approx [10/(\sqrt{E}) \oplus 2]$ %

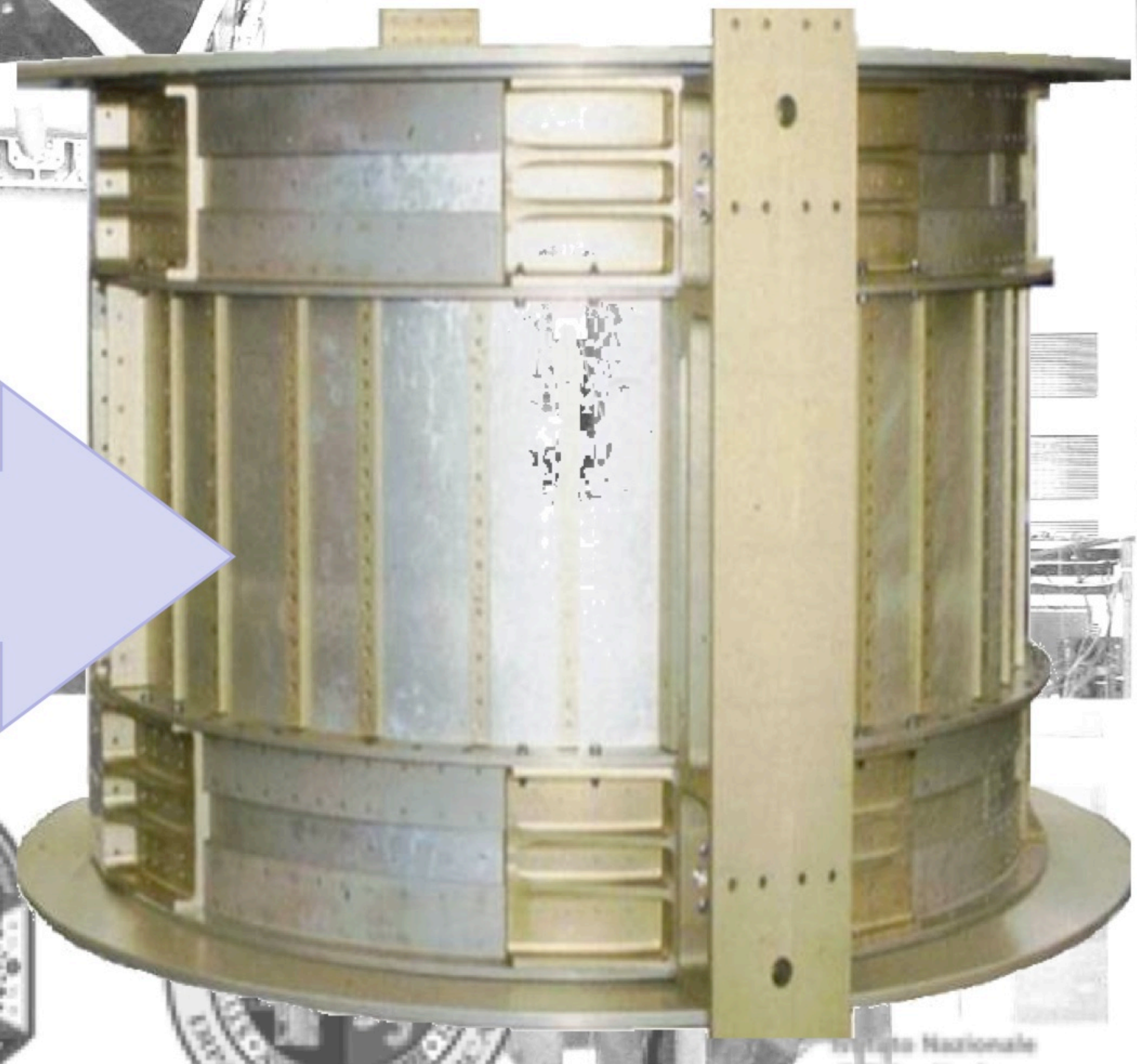
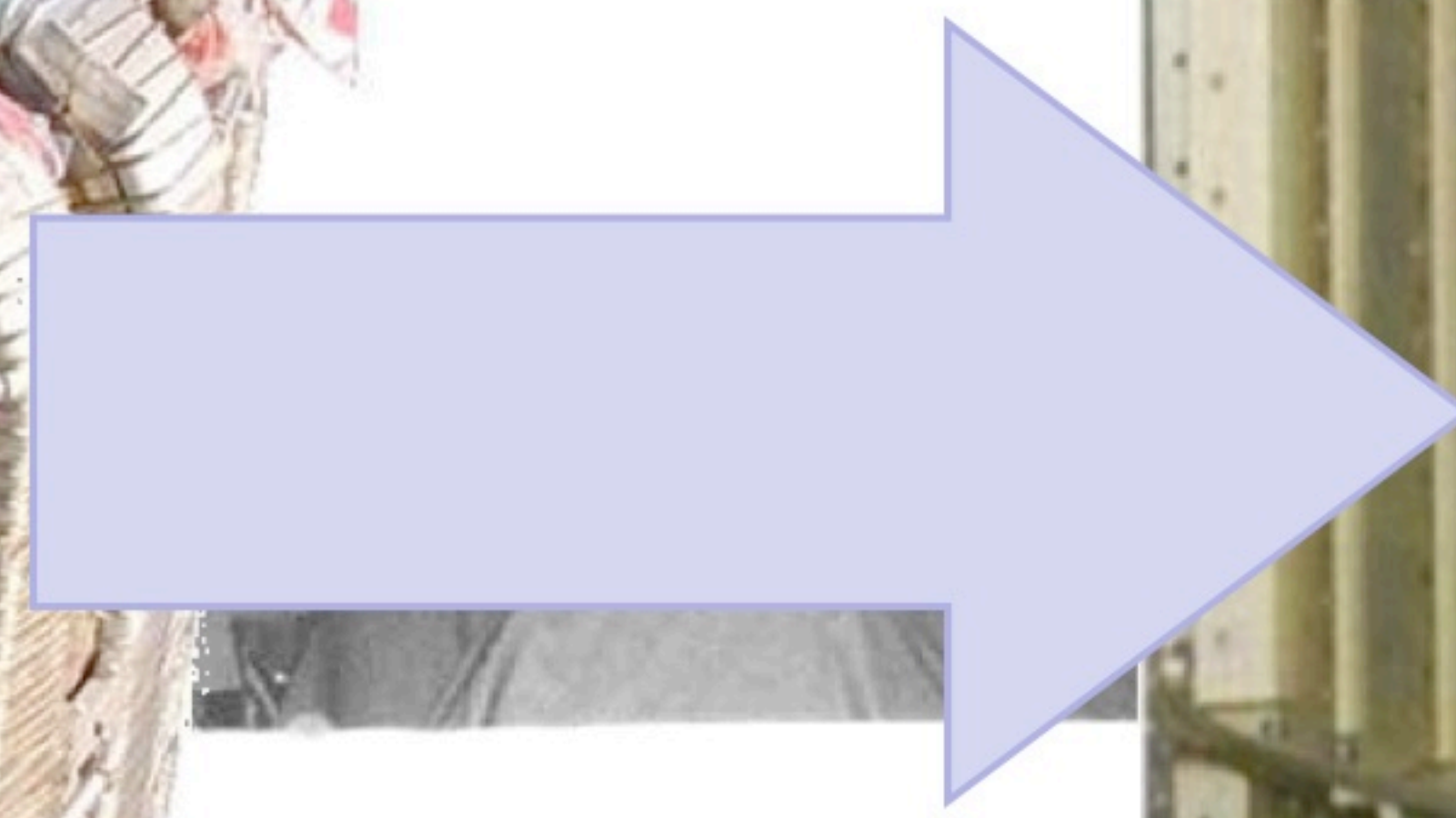
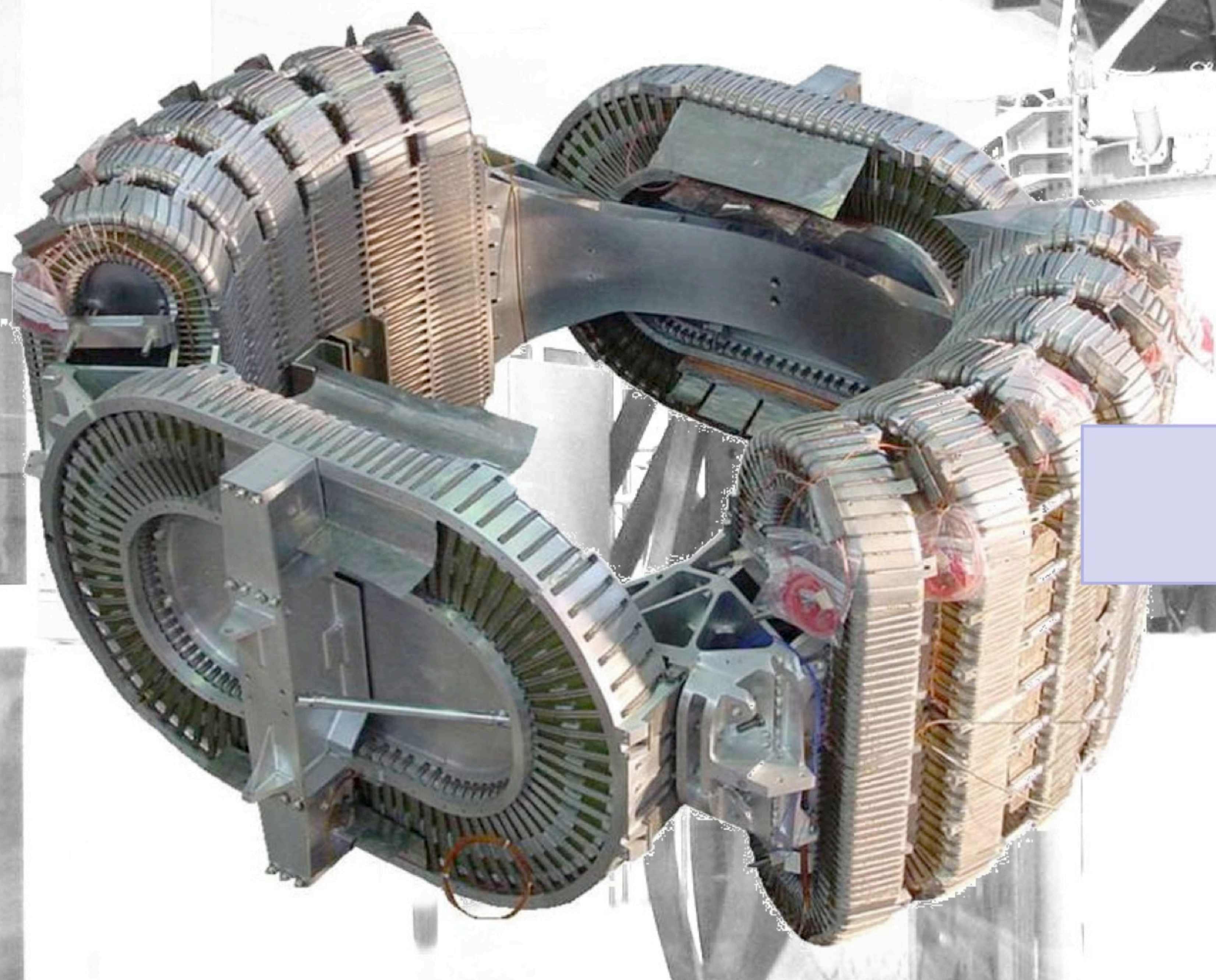
γ : ECAL (photon mode) or Tracker (conversion mode)



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AMS-02 The July Major Change: from SCM to PM

A very good reason: extension of the ISS program to 2020 (Feb 2010)!



Superconducting Magnet

$B = 0.87 \text{ T}$

Endurance = 3 years

Re-estimated as less than 3 years from ESTEC Thermo-Vacuum test (Apr 2010)

Permanent Magnet

$B = 0.13 \text{ T}$

Endurance $\approx \infty$

Easier to handle in space

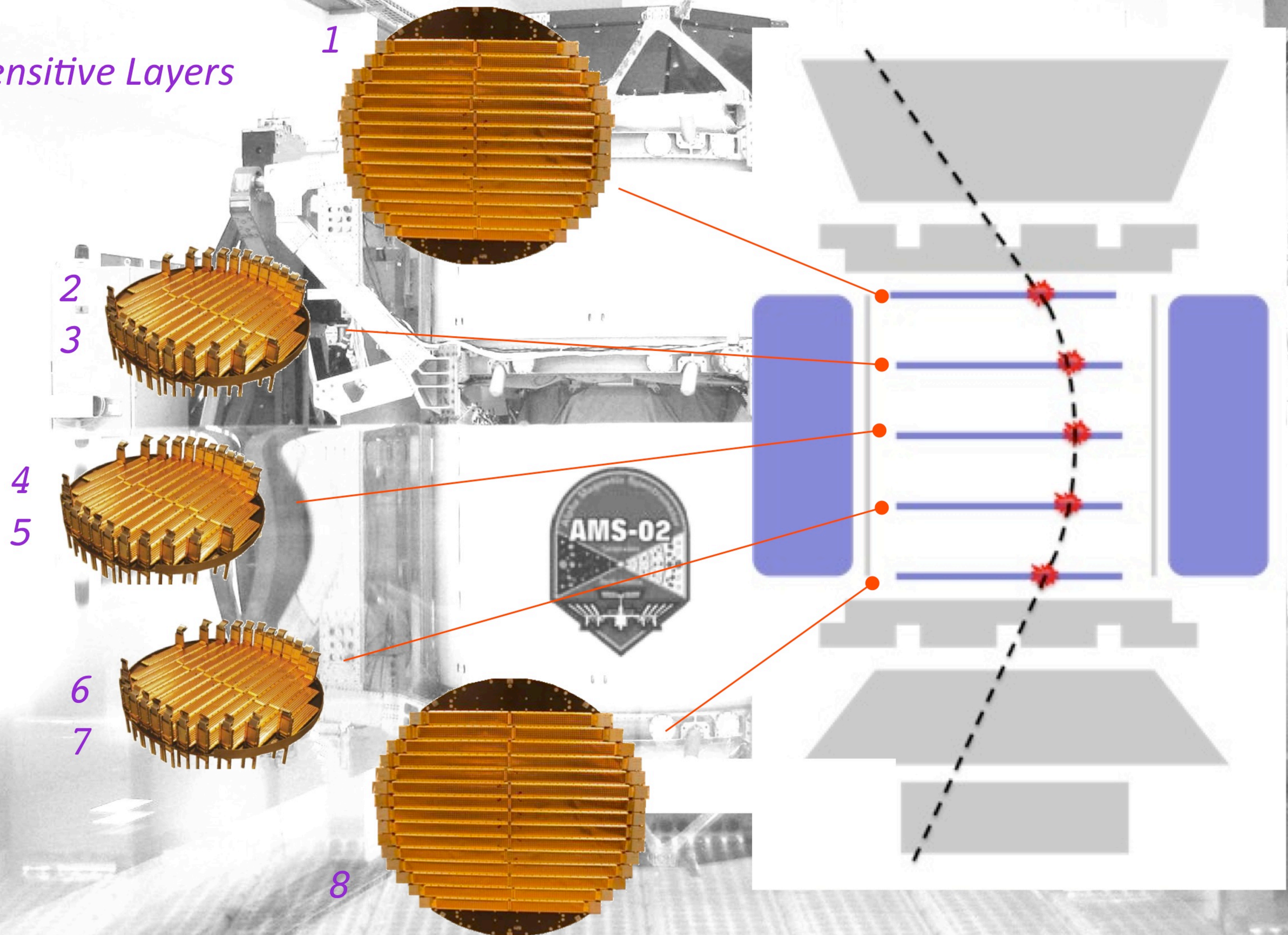
Already space-qualified

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But the designed momentum resolution would be affected → **Adjust the Tracker**

The AMS-02 Superconducting Magnet Configuration

Tracker Sensitive Layers

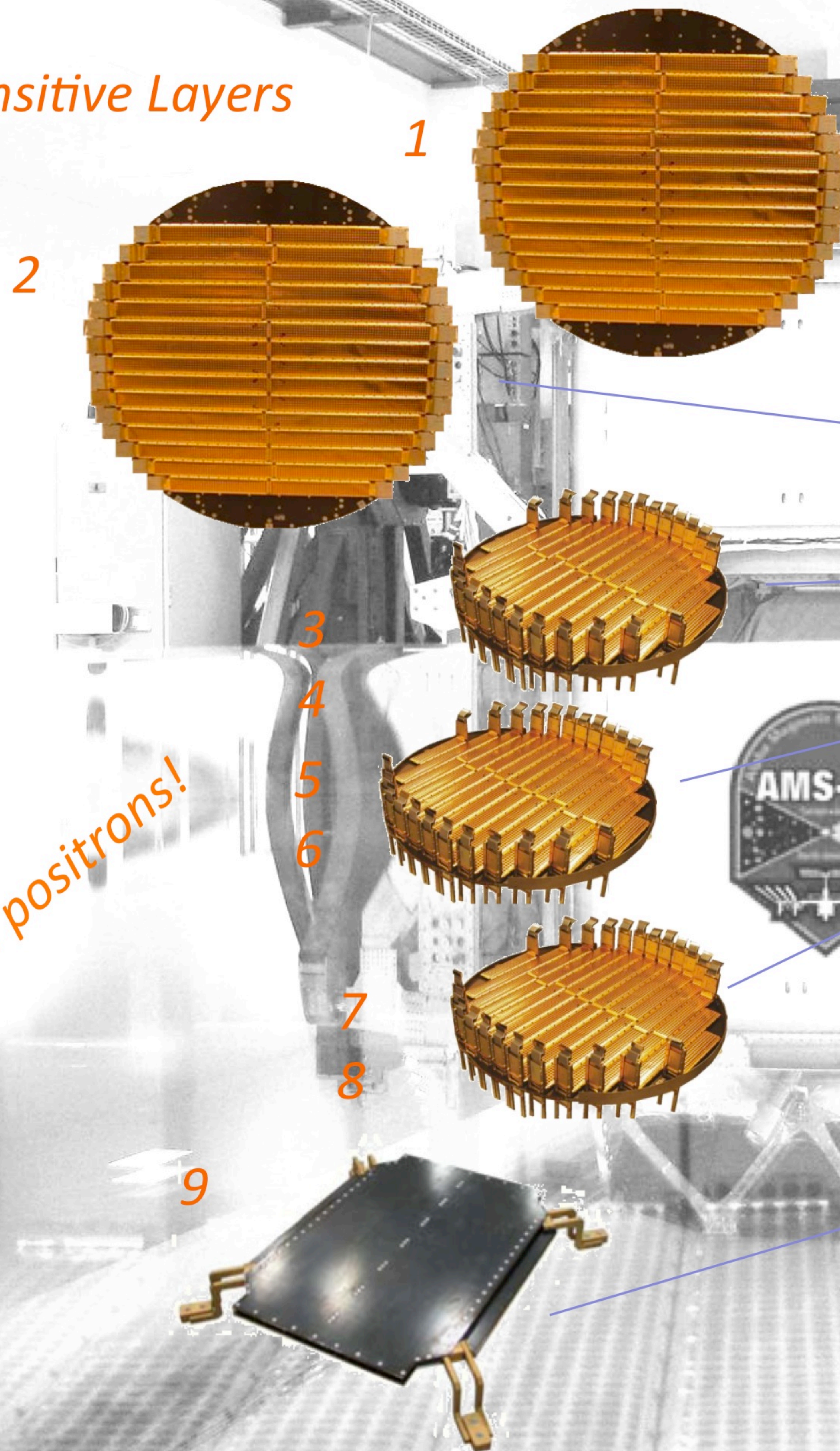


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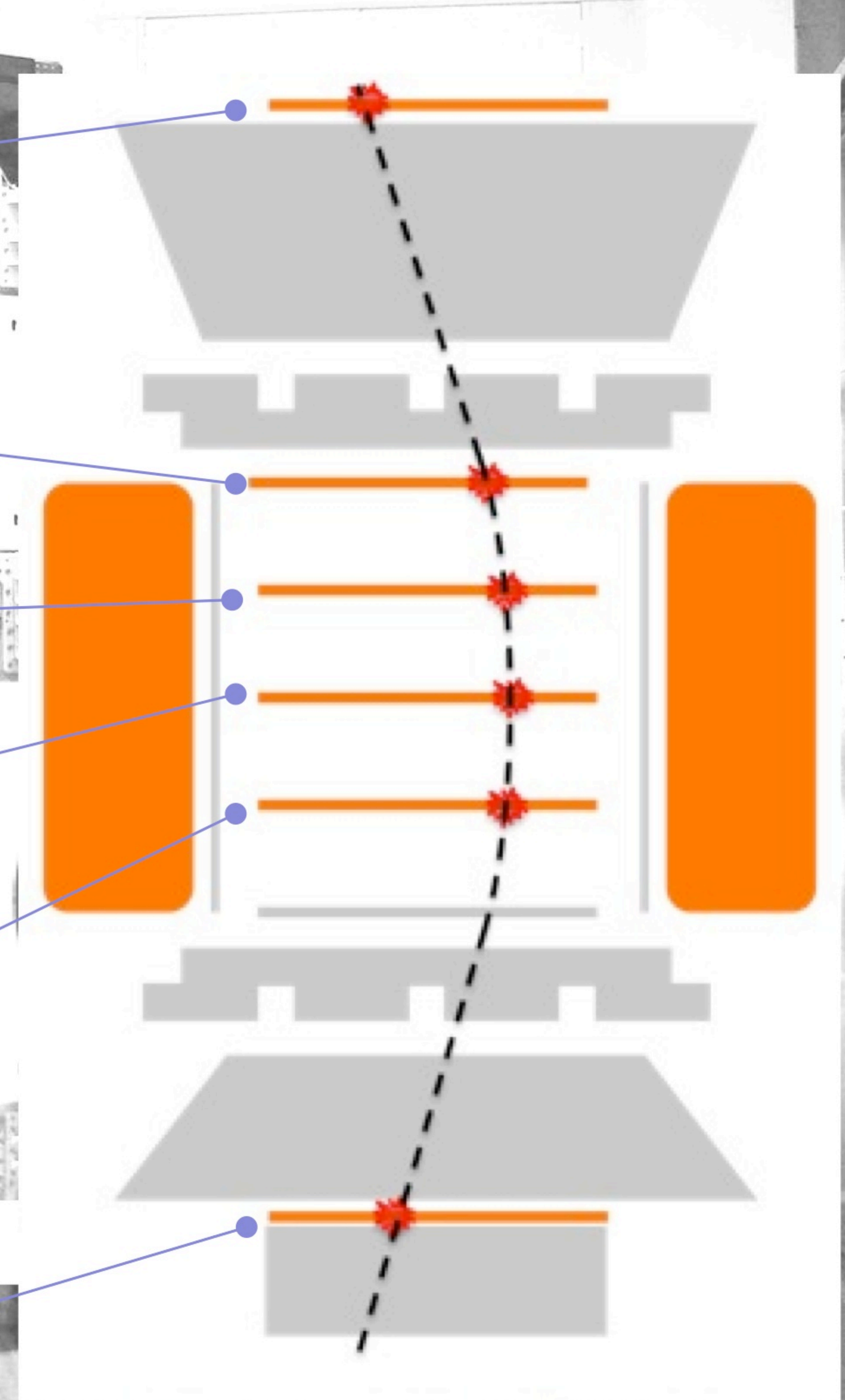
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The AMS-02 Permanent Magnet Configuration

Tracker Sensitive Layers



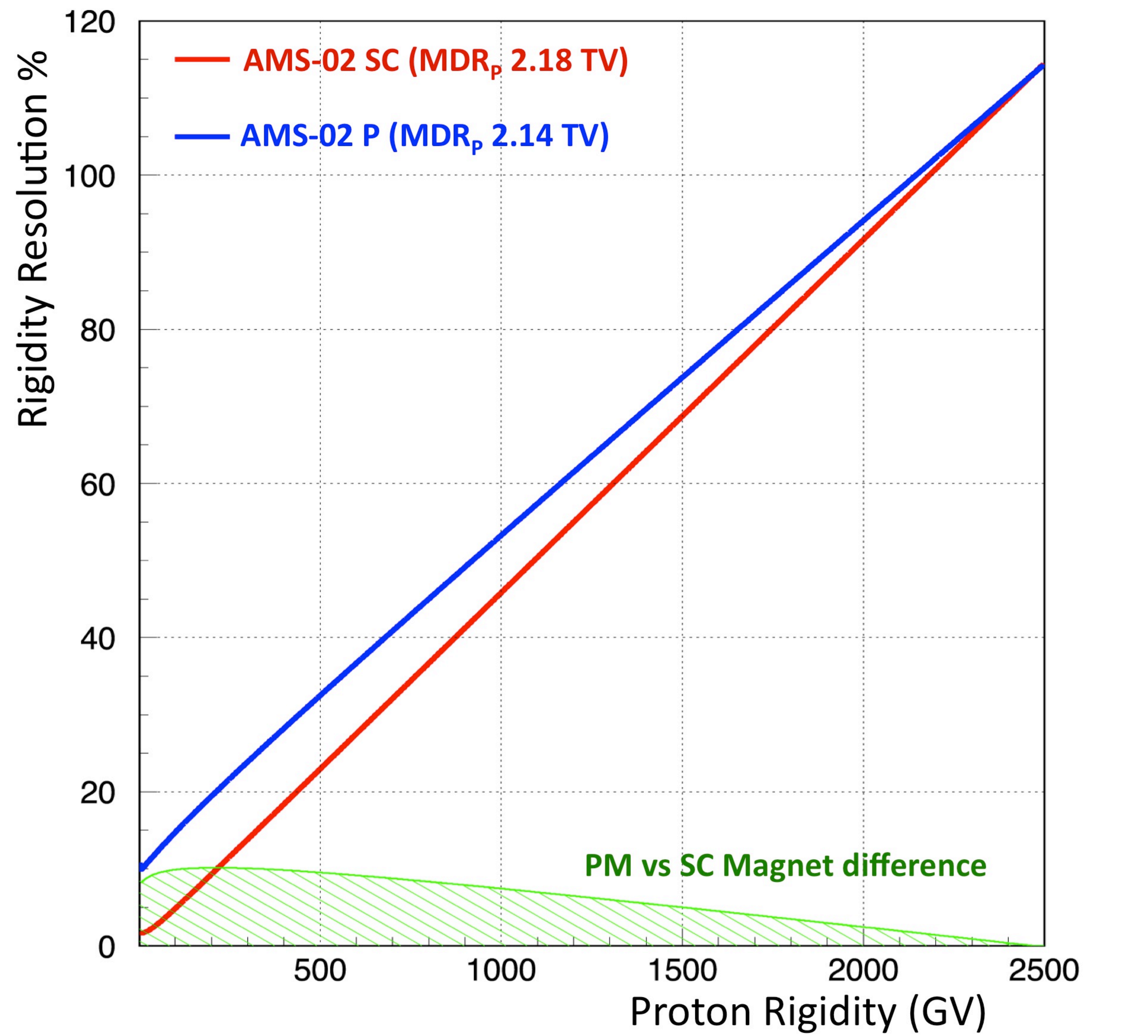
Optimized for positrons!



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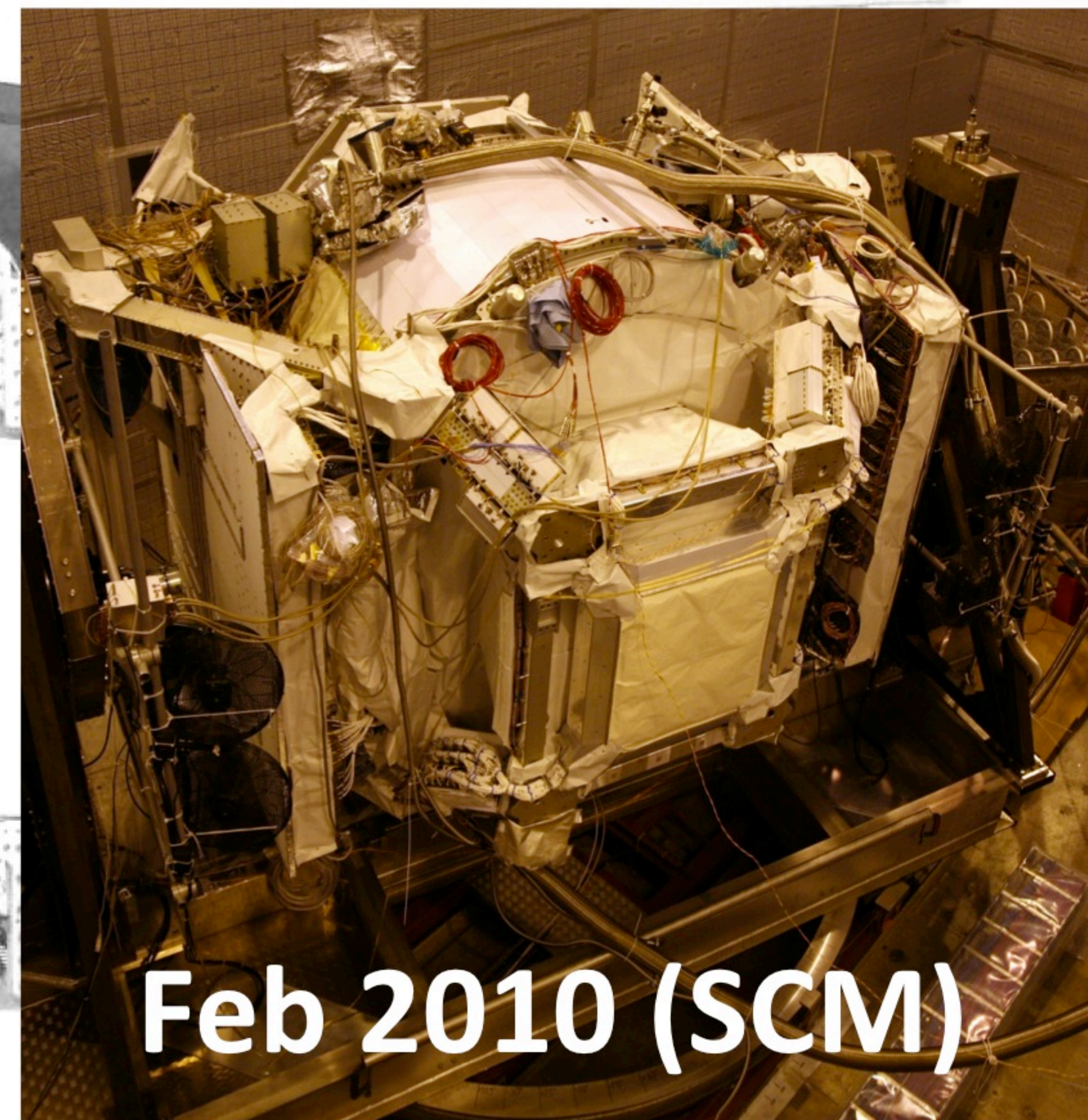
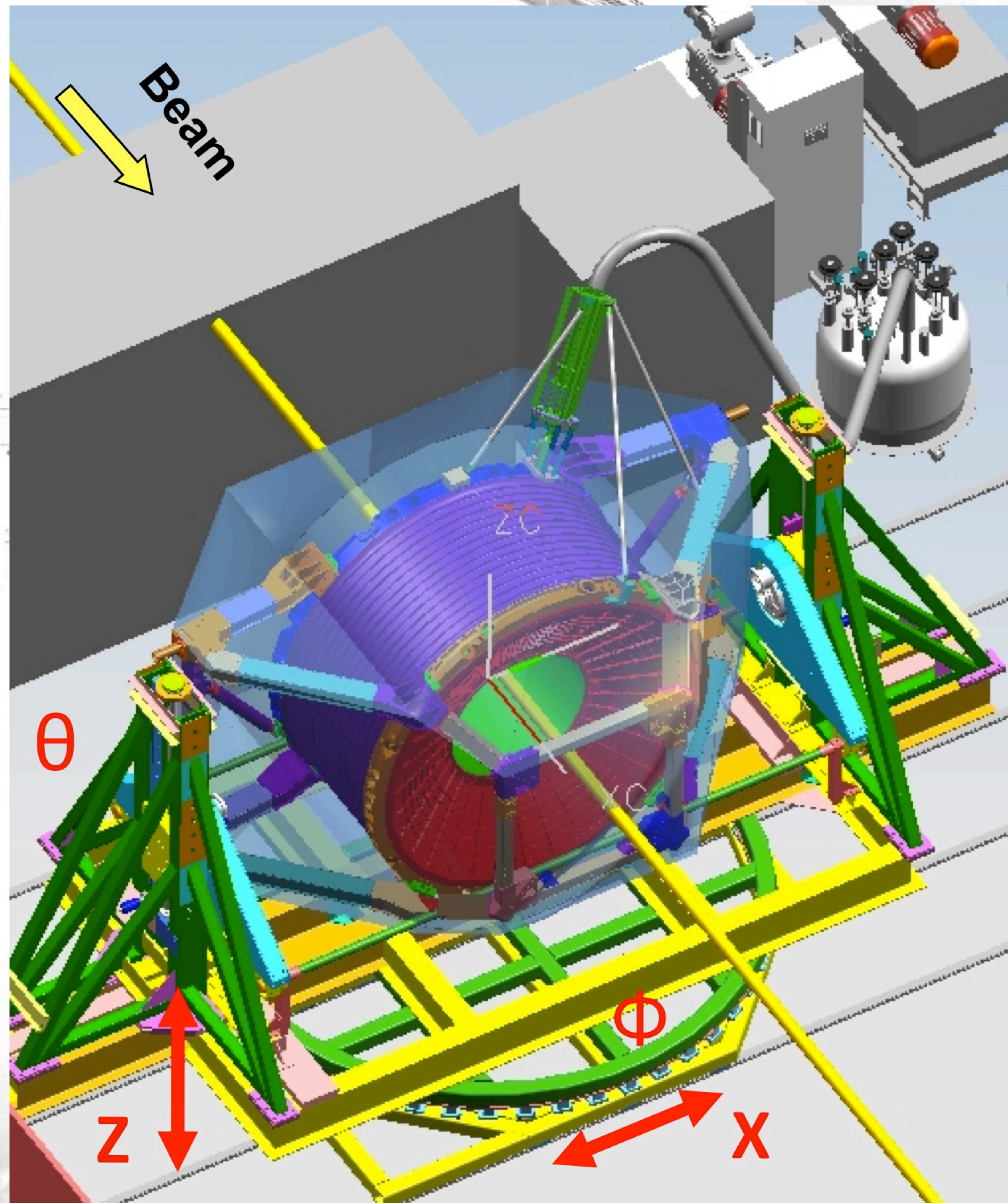
The Momentum Resolution: MonteCarlo



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The AMS-02 Test Beams



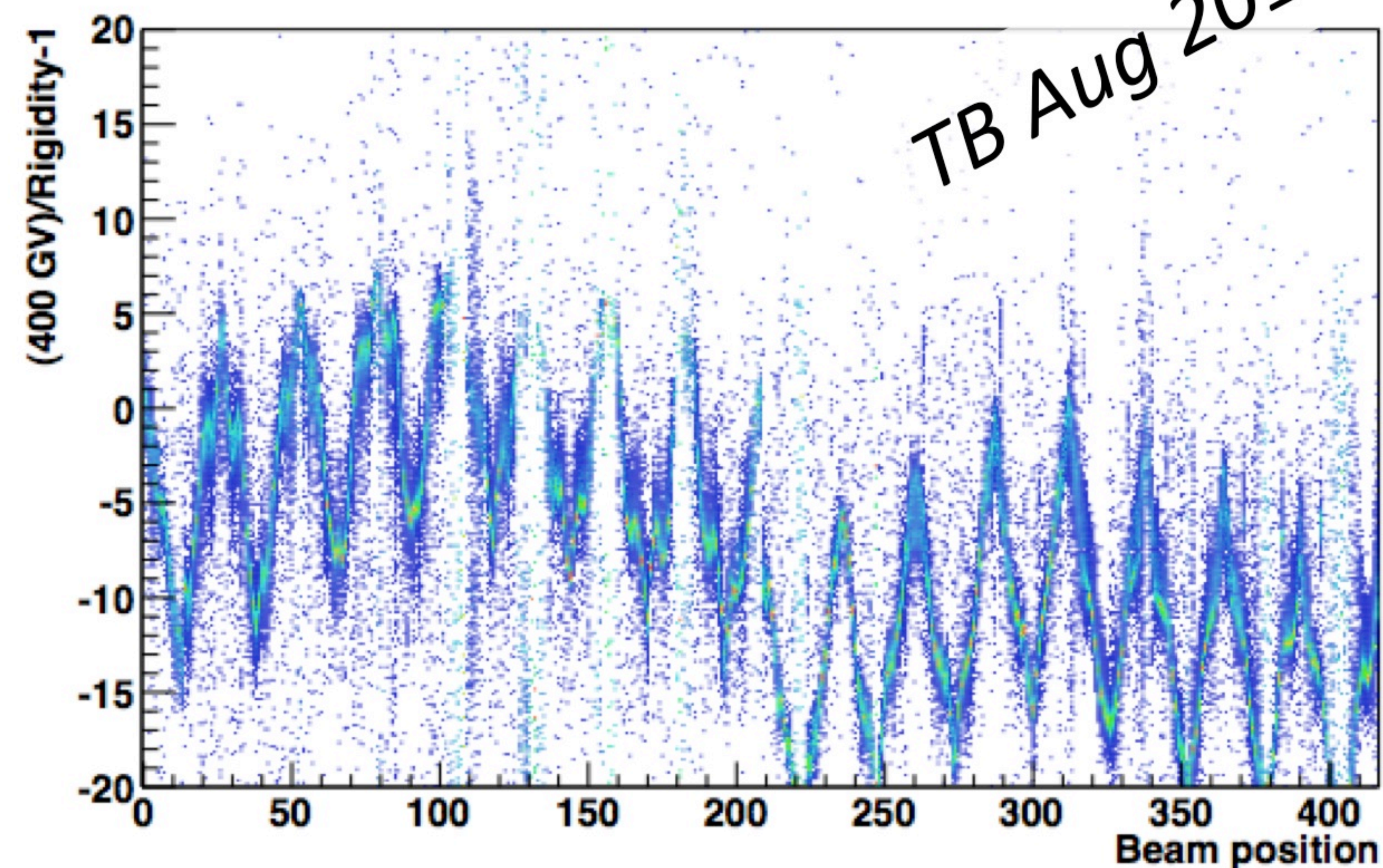
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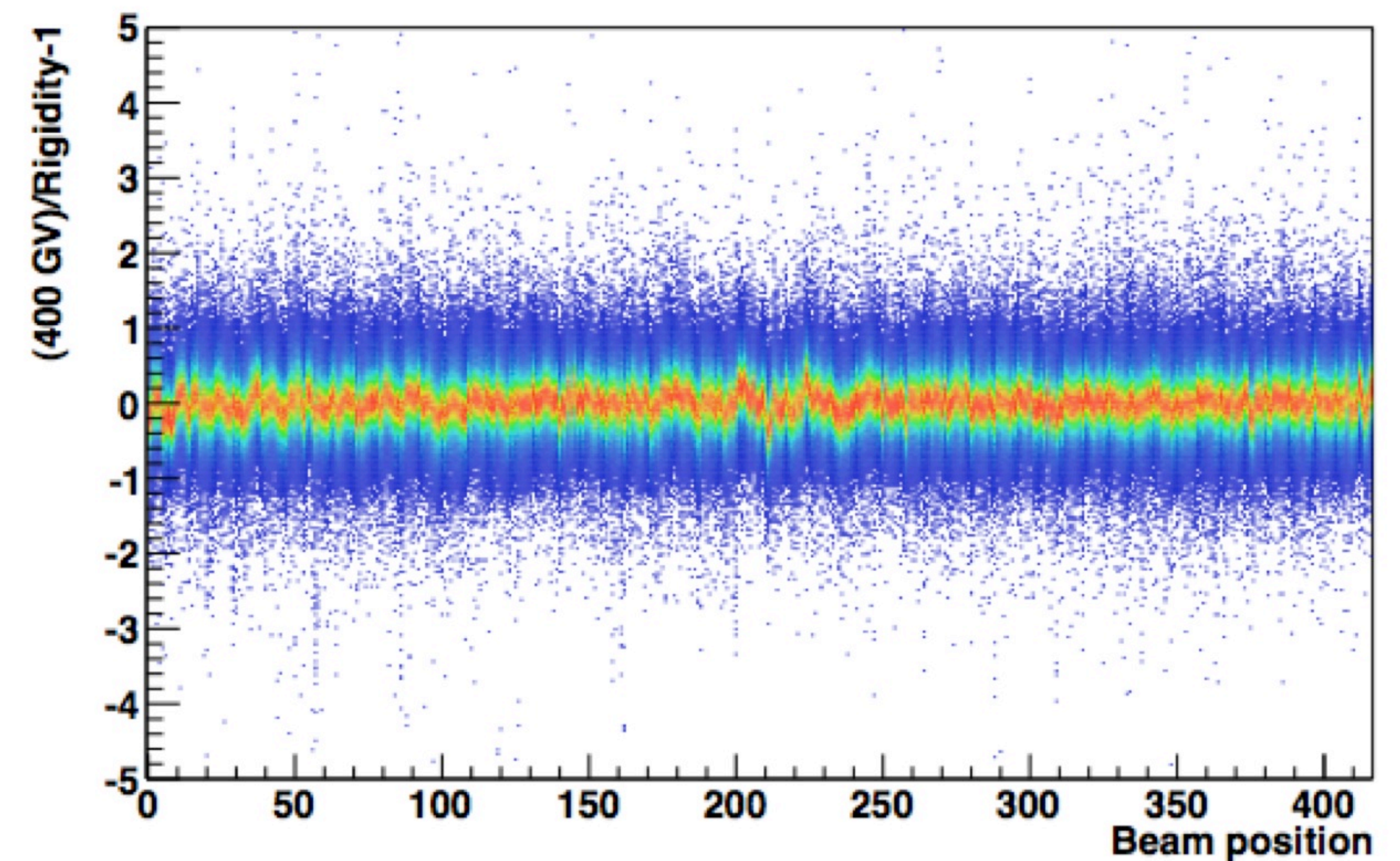
The AMS-02 Test Beam Purposes

- > Tracker Alignment at Fixed Rigidity (p at 400 GV/c)
- > Alignment Check with Positive and Negative Samples (π^\pm and e^\pm at 180, 290, ... GV/c)
- > Tracker Resolution and Charge Confusion
- > e/p Separation Studies (ECAL, TRD)
- > Sub-detectors Threshold and Uniformity Studies

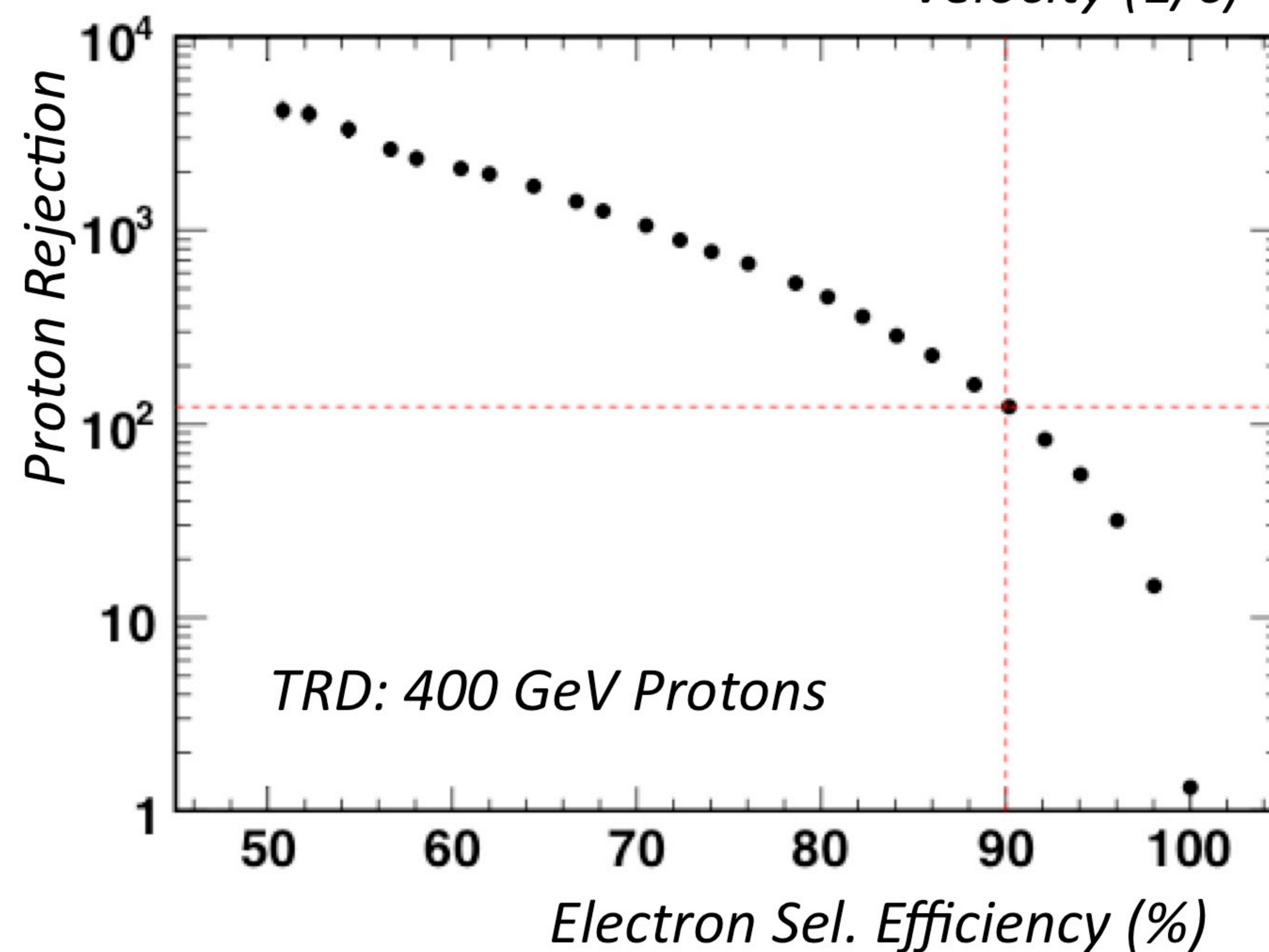
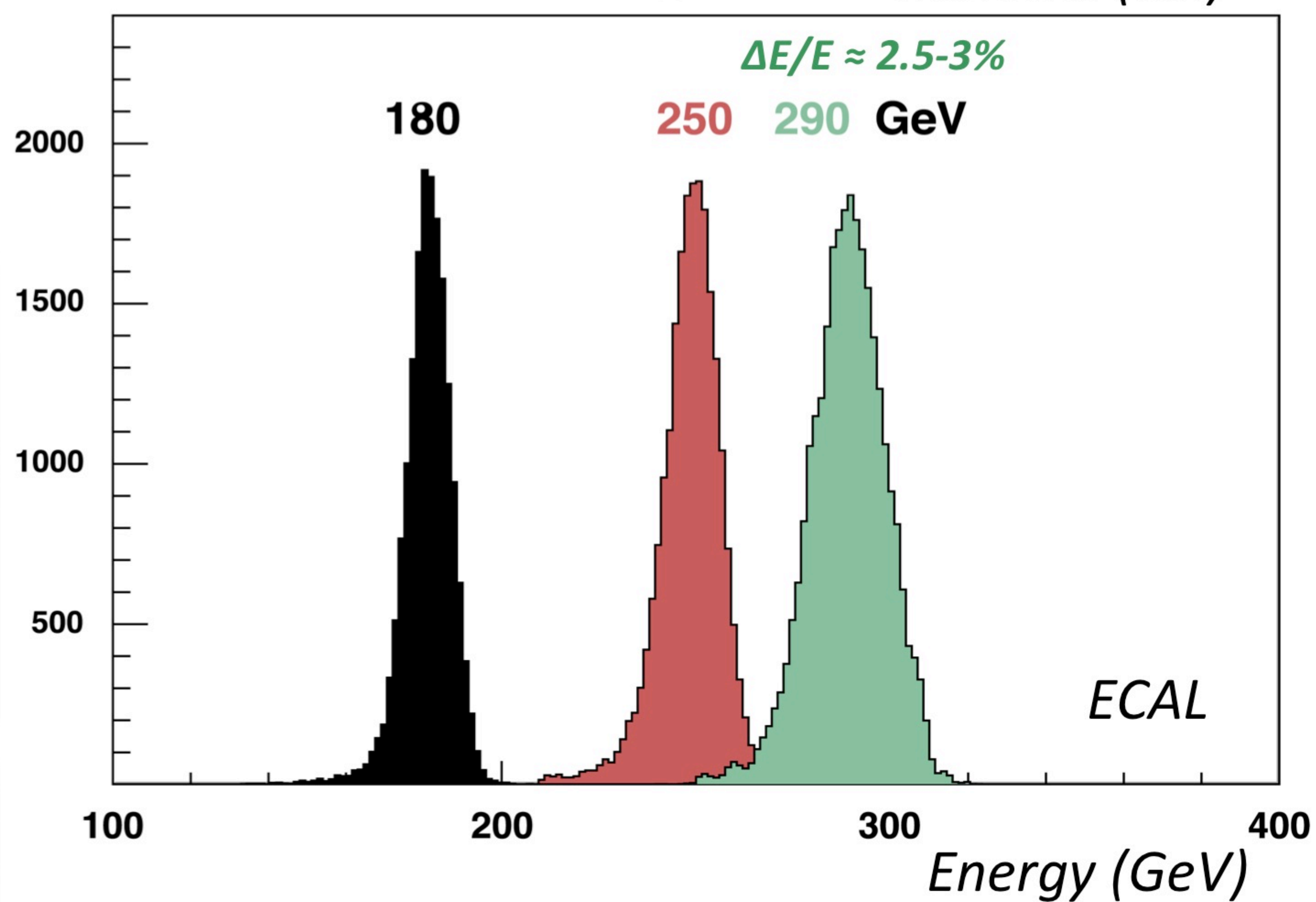
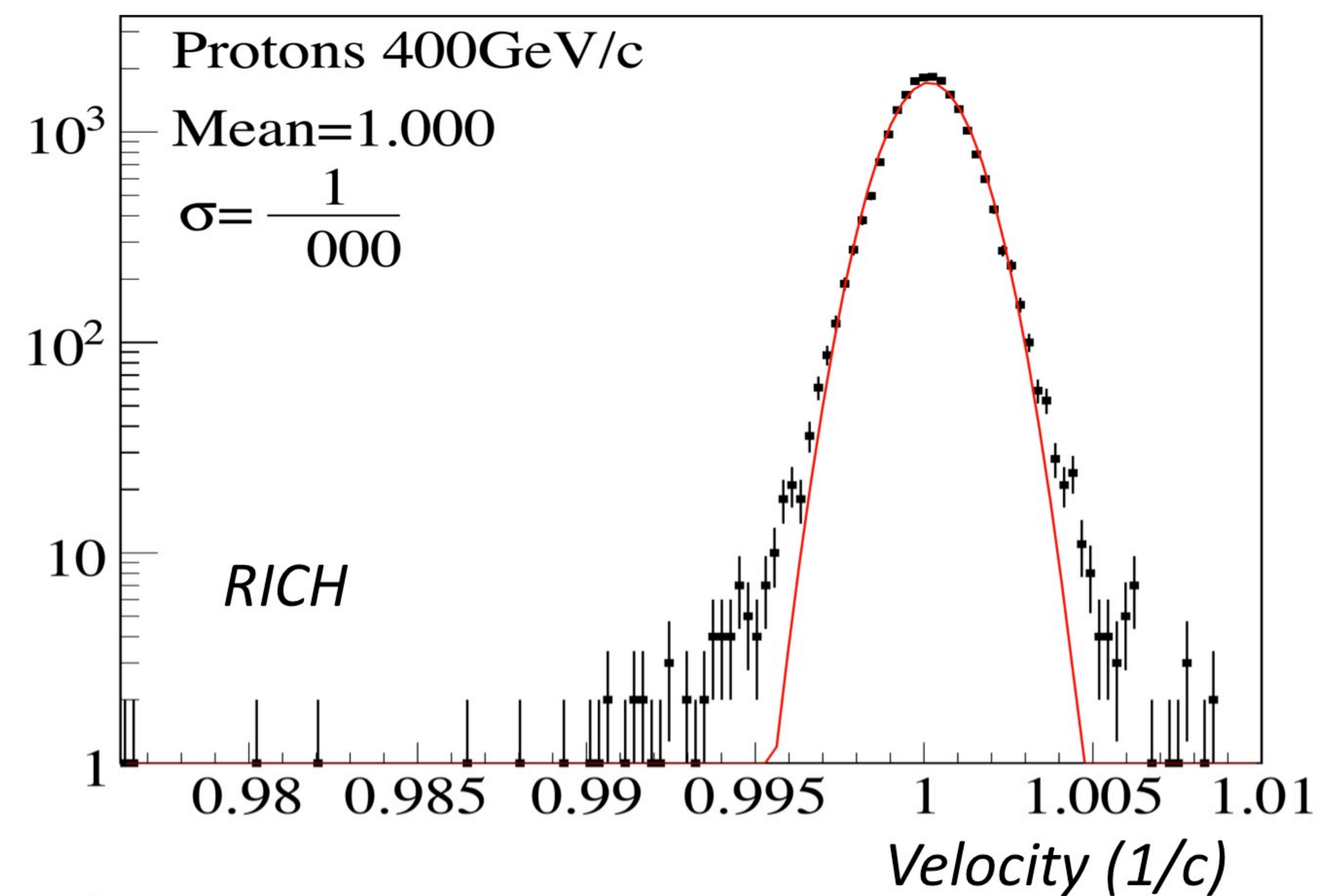
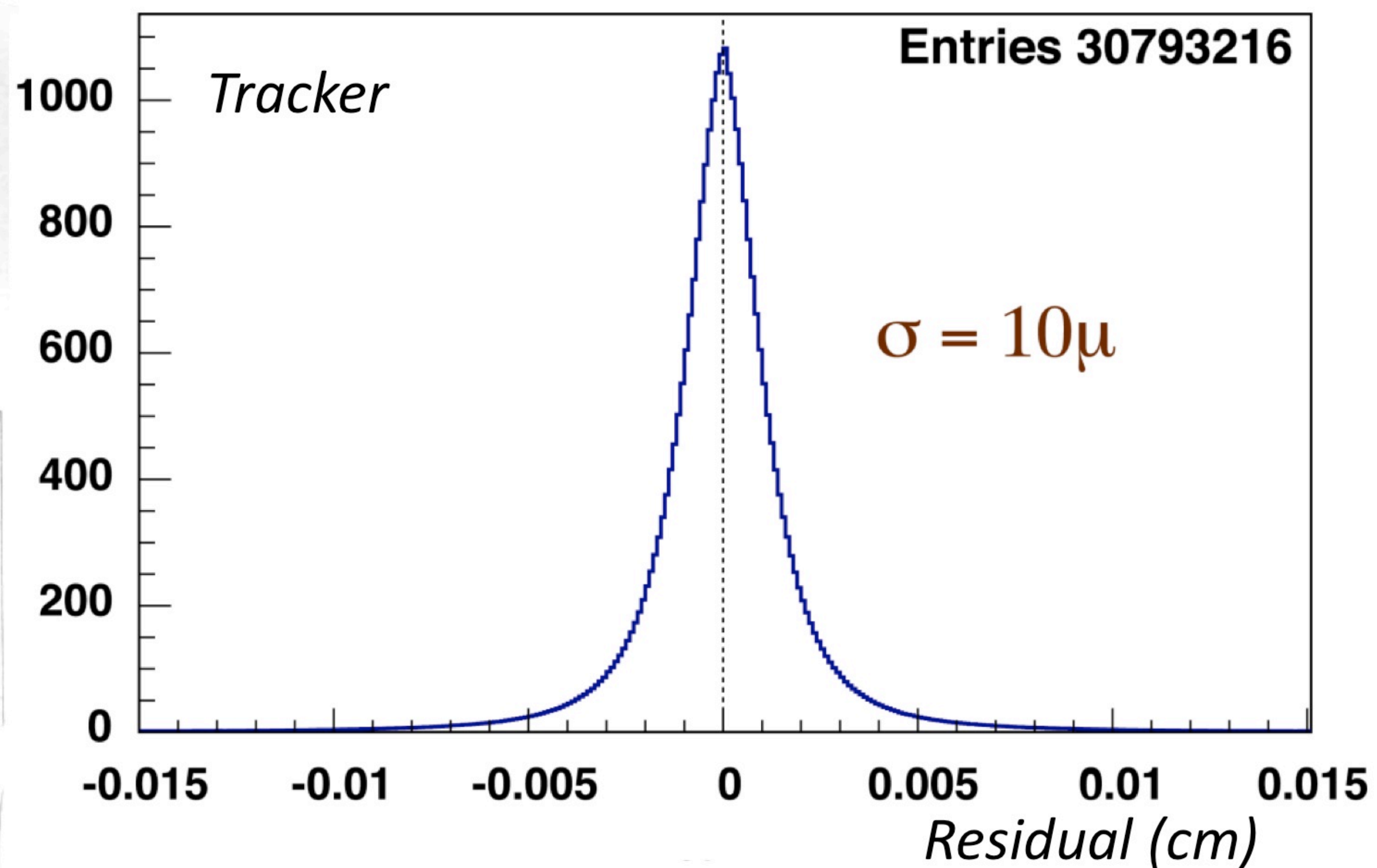
Before alignment



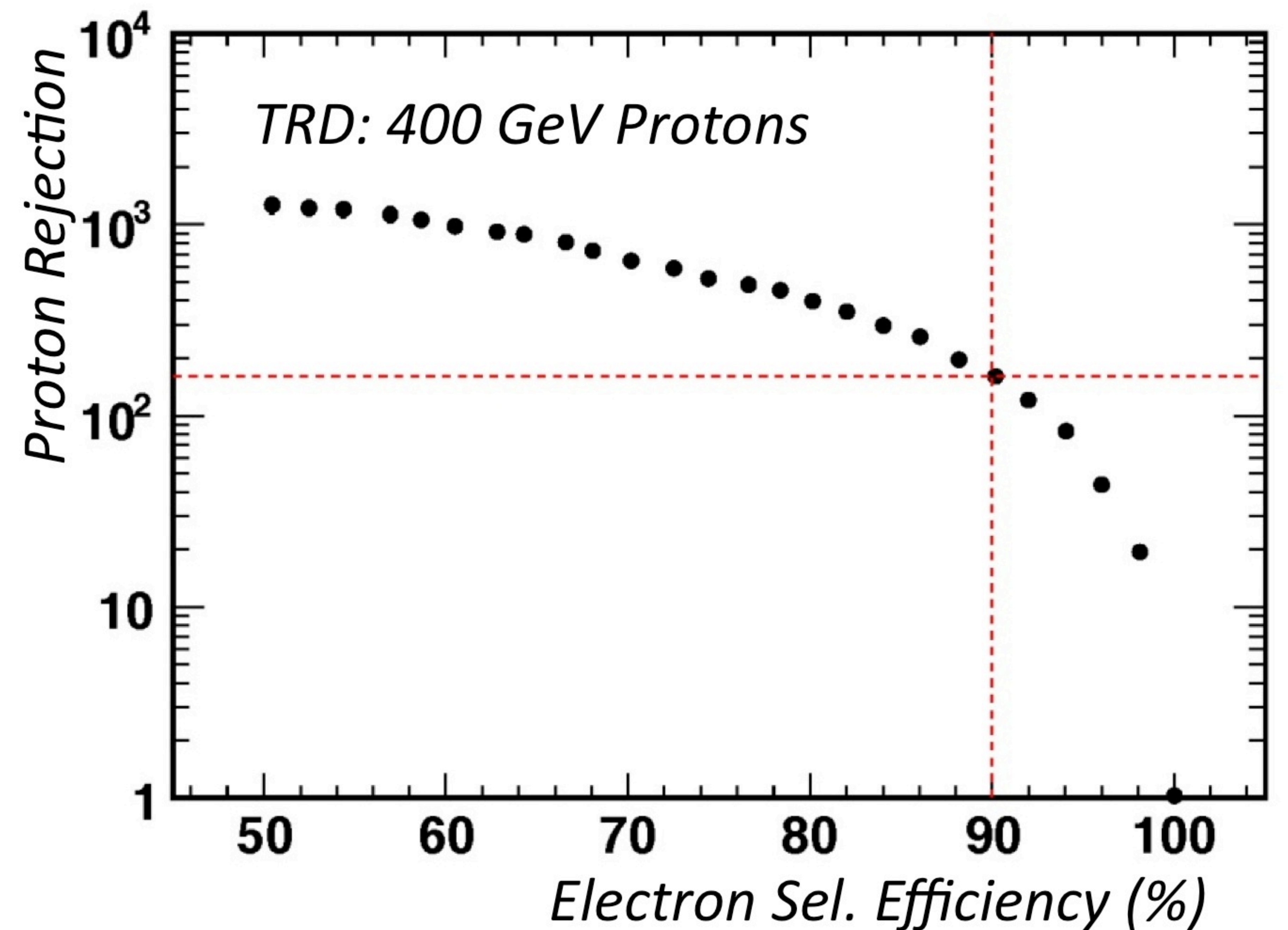
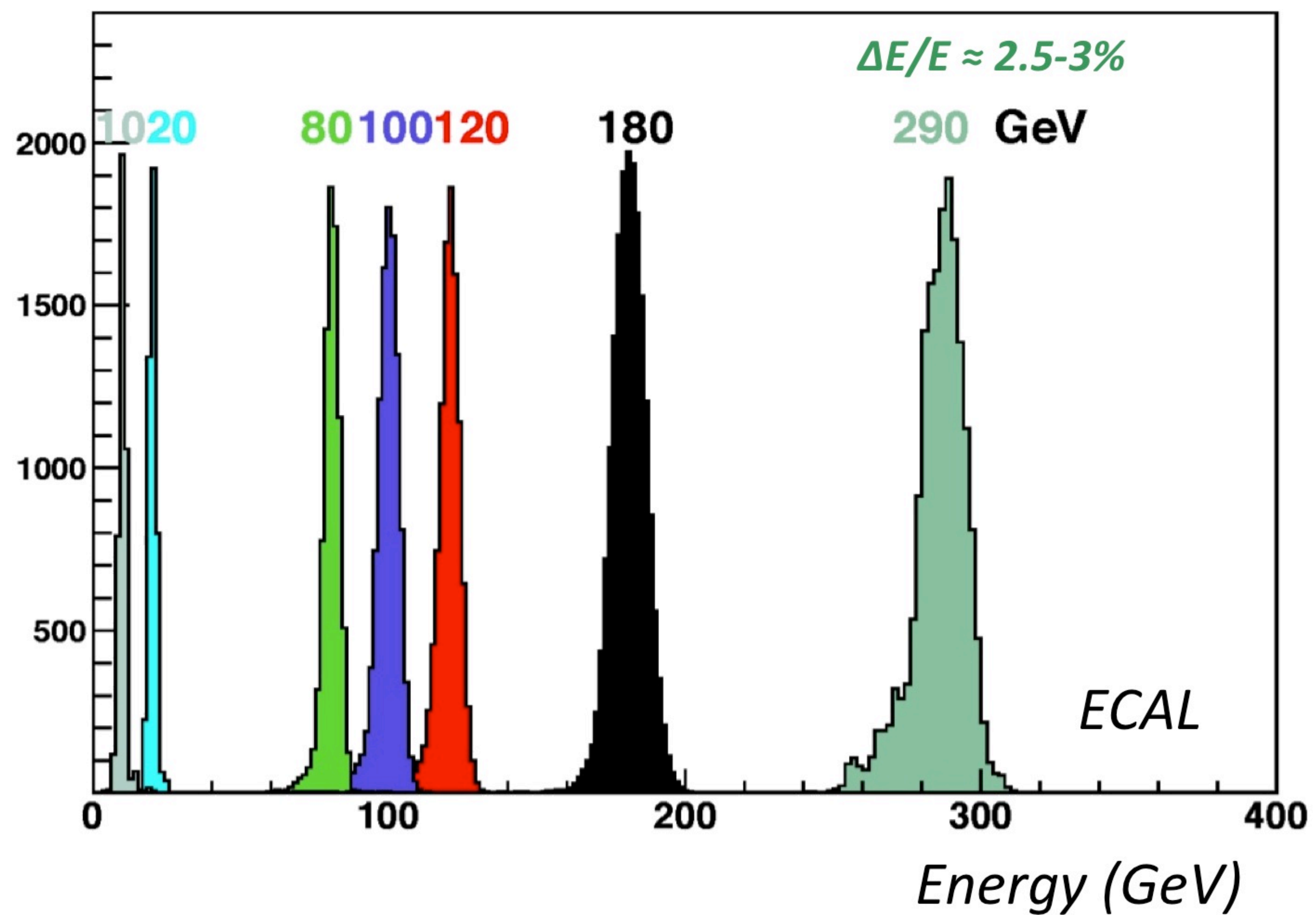
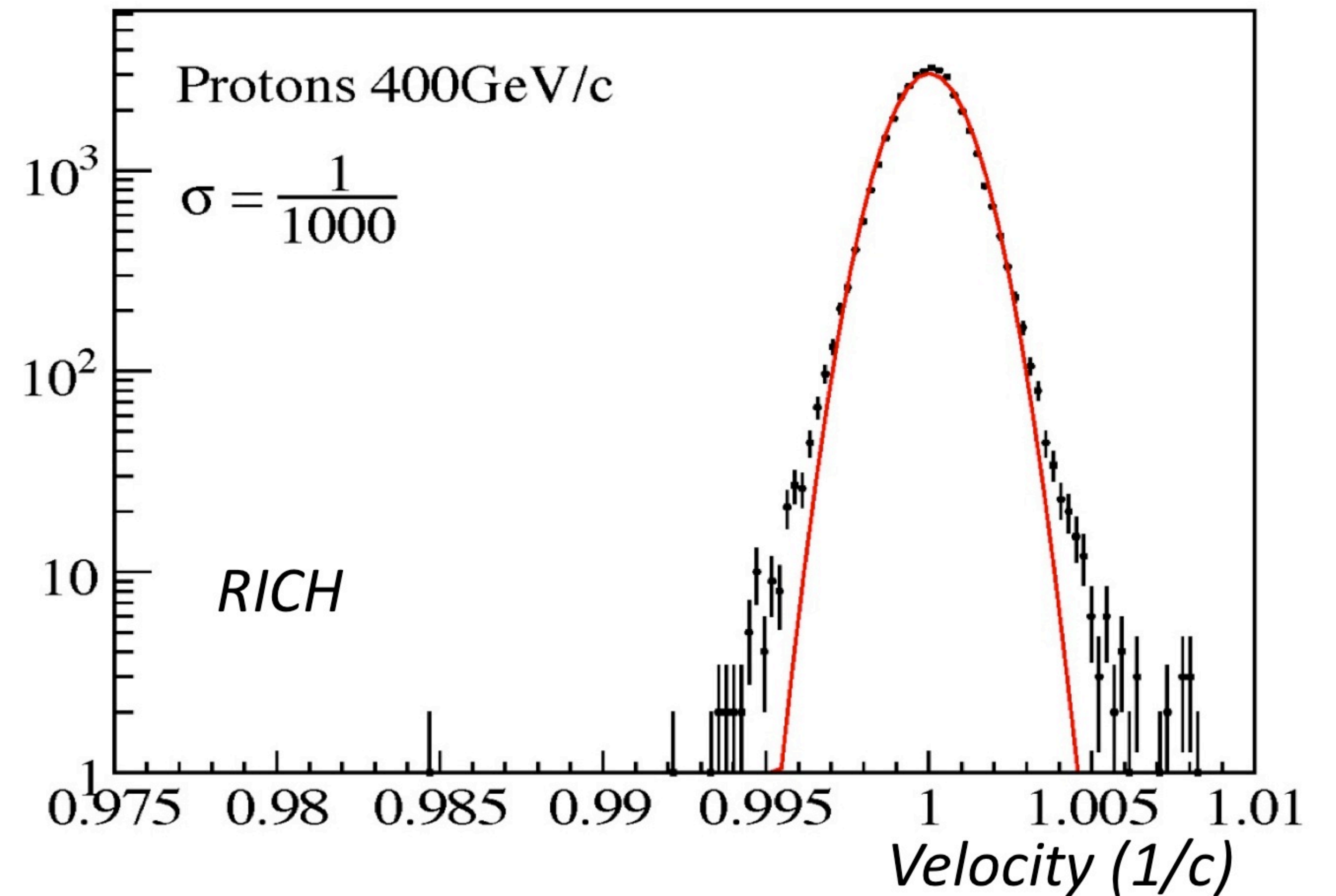
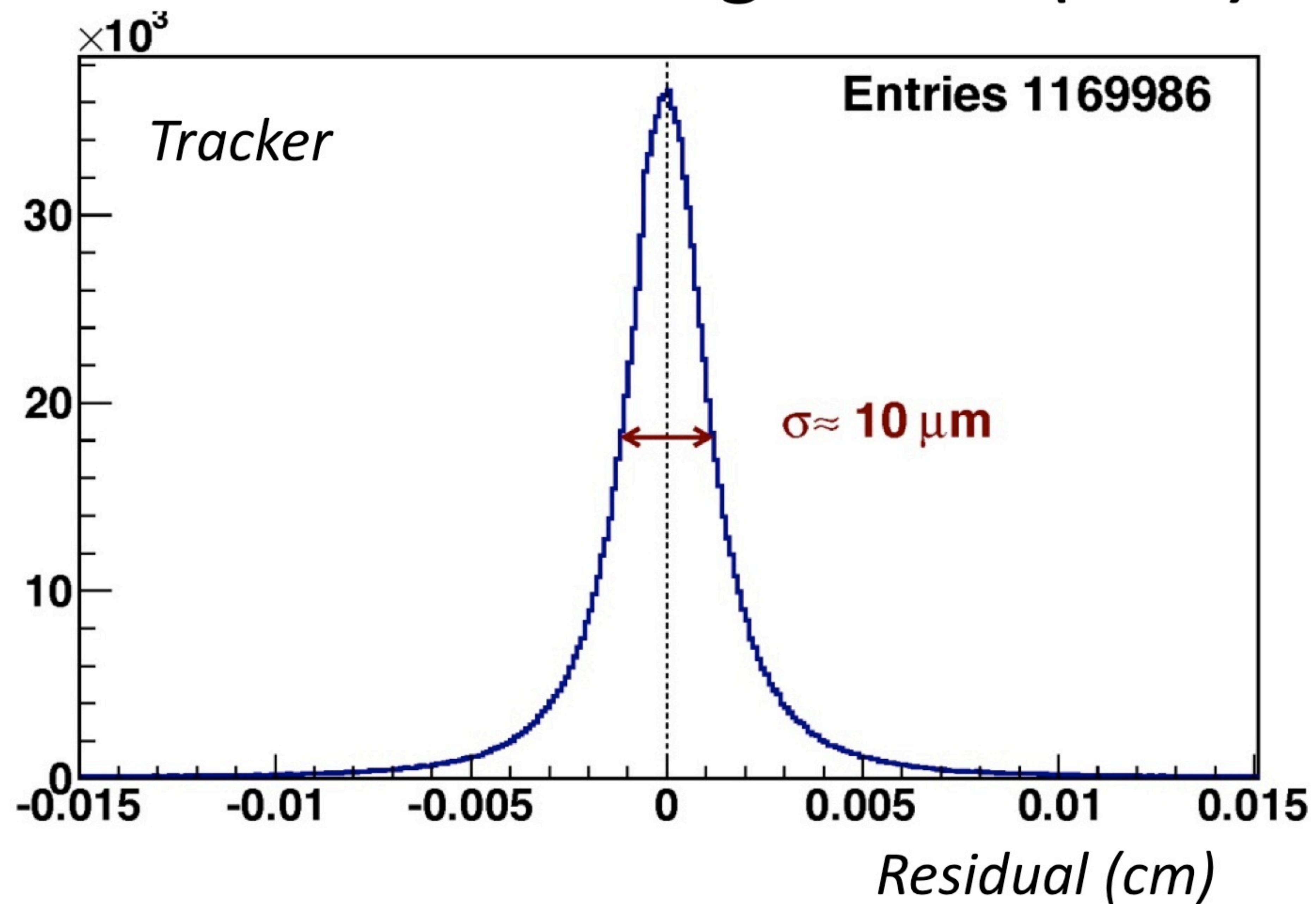
After alignment



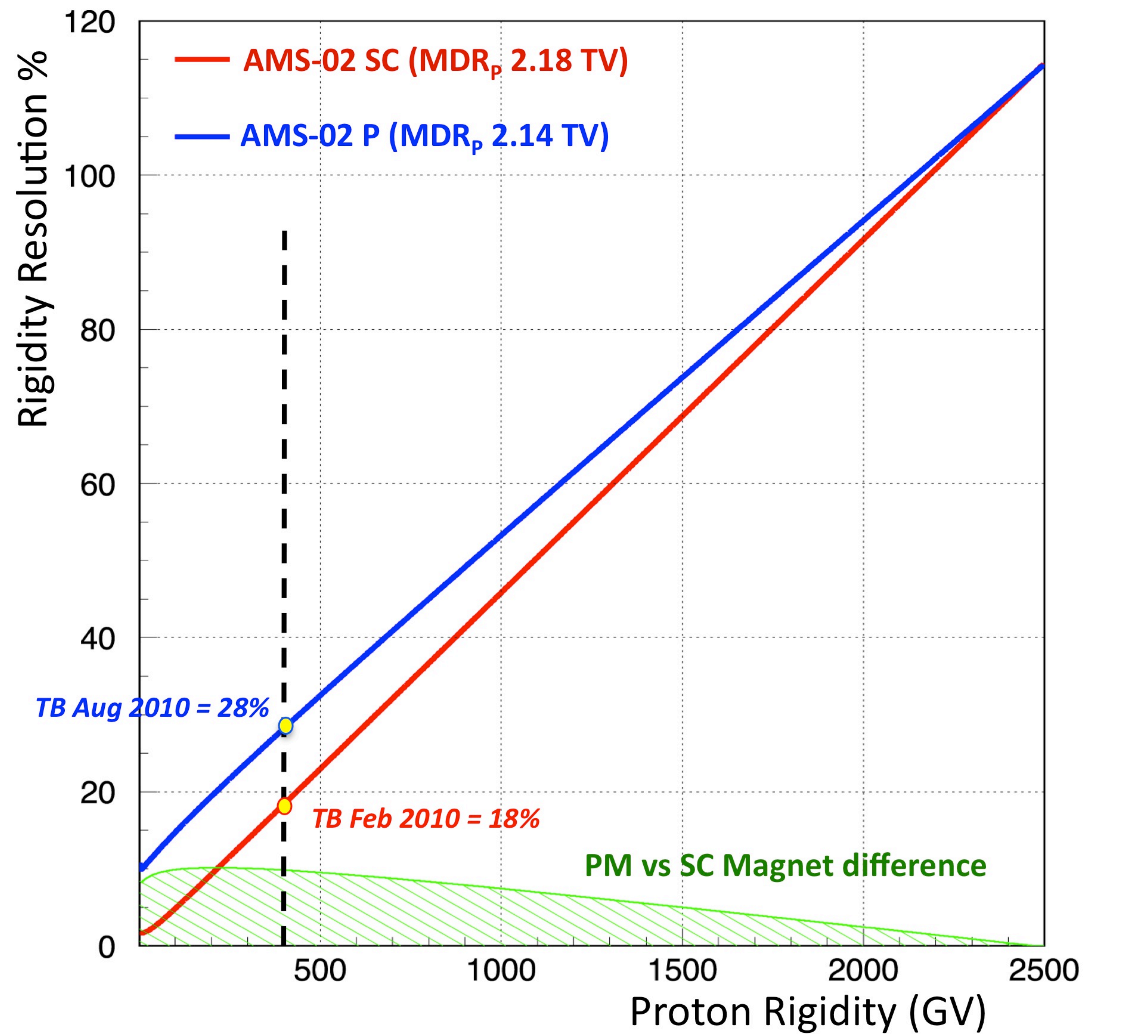
Test Beam Feb 2010 (SCM) Results



Test Beam Aug 2010 (PM) Results



The Momentum Resolution Estimation from TBs



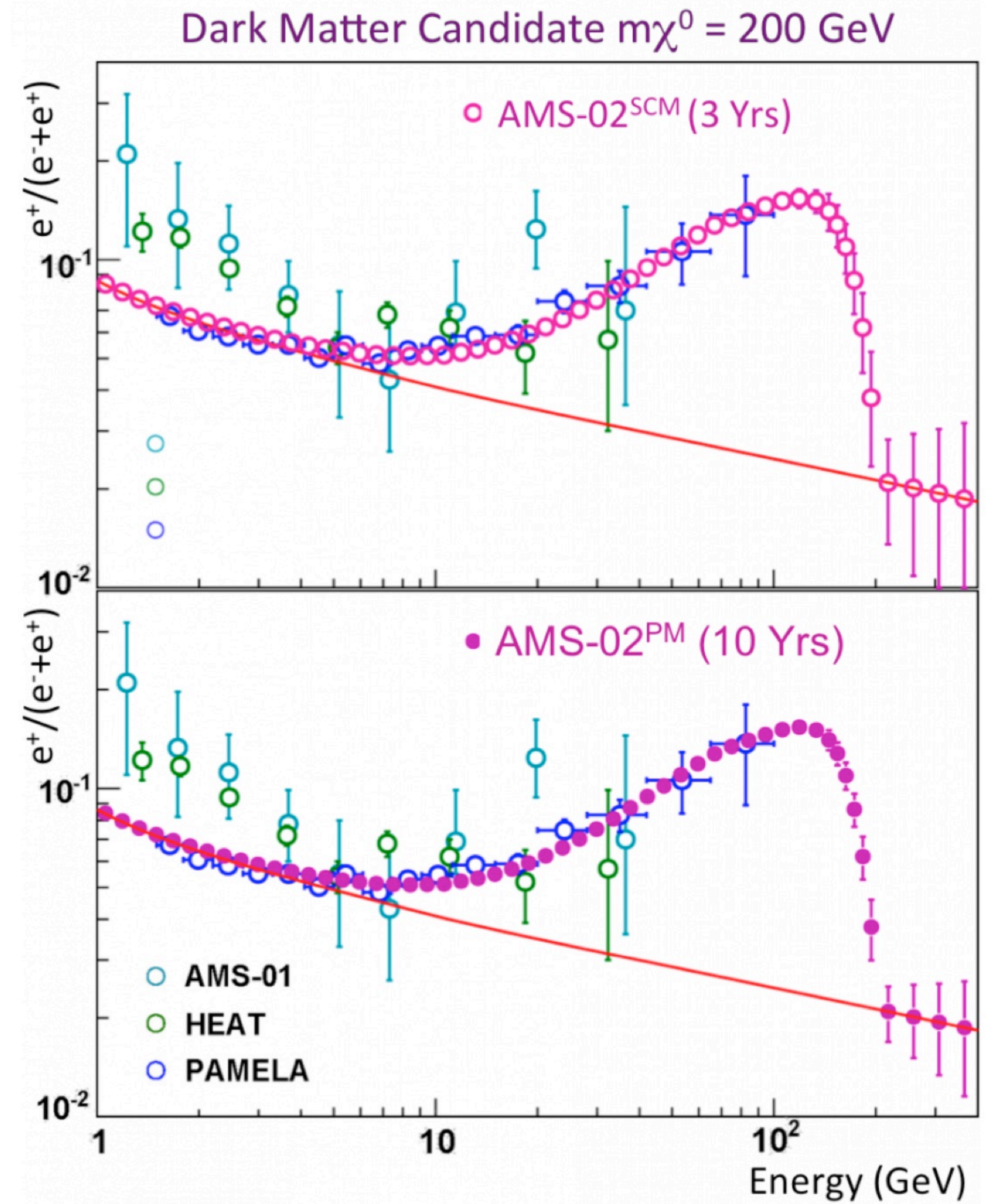
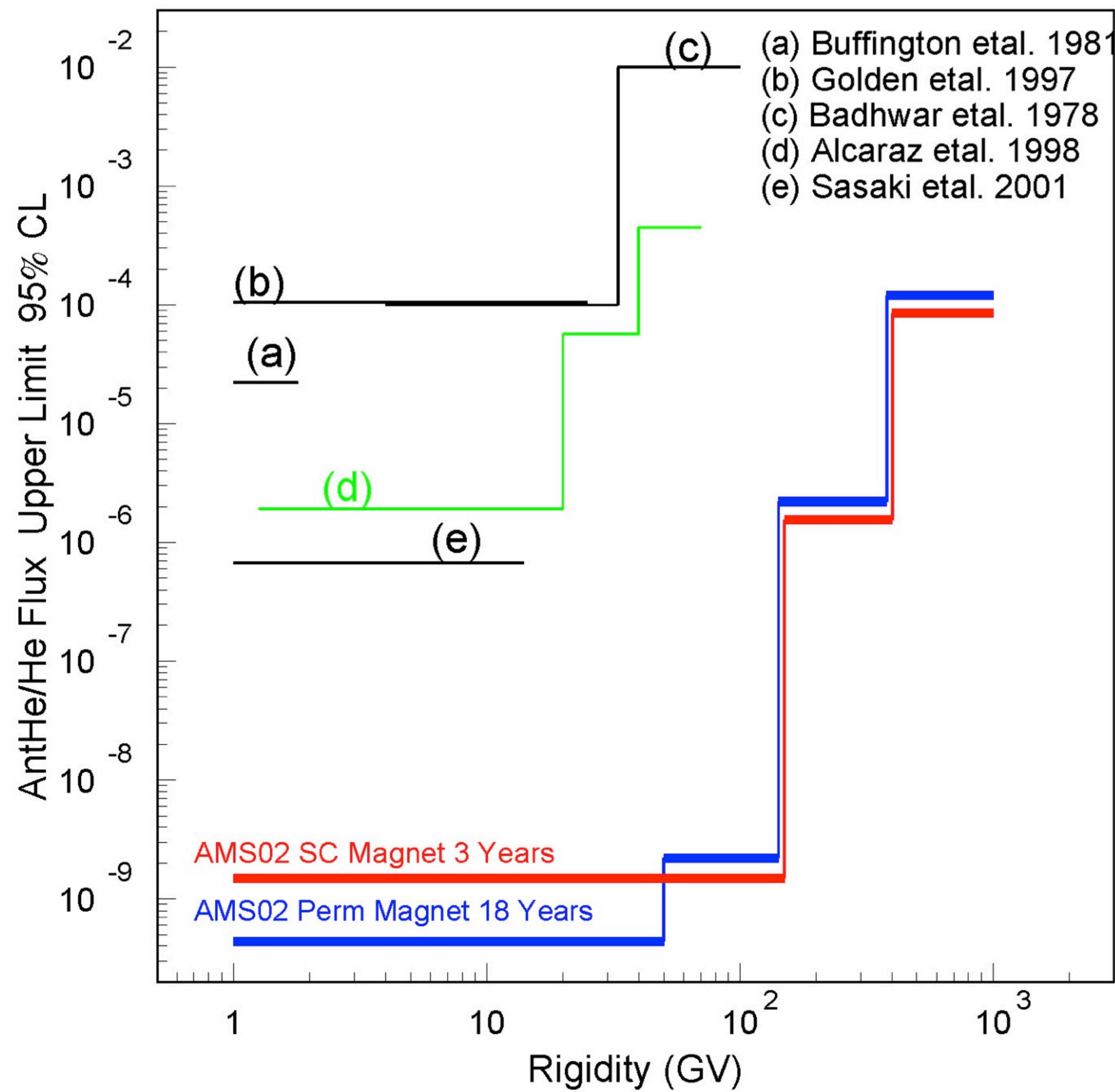
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Expected Performances

Antimatter

Dark Matter "Best" Channel



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Conclusions

AMS-02@ SSPF (KSC)



- > AMS-02 PM is fully tested and validated
- > AMS-02 PM is ready at KSC
- > AMS-02 will be launched on Feb 2011
- > Follow AMS-02 adventures on www.ams02.org

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