GALPROP WebRun

http://www

galprop.stanford.edv/

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for the GALPROP development team

What's new in v.54

- Shared-memory parallel support with OpenMP: multi-processor machines
- Memory usage optimization
- HEALPix output of γ-ray and synchrotron skymaps
- MapCube output for compatibility with Fermi-LAT Science Tools software
- γ-ray skymaps output in Galactocentric rings to facilitate spatial analysis of the Galactic diffuse γ-ray emission
- More accurate line-of-sight integration for computing diffuse emission skymaps
- 3D modeling of the Galactic magnetic field, both regular and random components
- Calculations of synchrotron skymaps, using both regular and random magnetic fields
- New improved gas maps, which are computed using recent H I and CO surveys
- A new calculation of the Galactic interstellar radiation field using the FRaNKIE code
- Increased efficiency of anisotropic inverse Compton scattering calculations
- GALPROP code is compiled to a library for easy linking with other codes (e.g., DarkSUSY, SuperBayeS
- Improved configuration management via the GNU autotools. Multiple *NIX system and compiler targets (gcc, intel, llvm, open64) are supported

GALPROP WebRun

- GALPROP WebRun is a service that allows to run GALPROP via the WWW
- No local installation of the code or related libraries is necessary, only a web browser is required
 - Available at http://galprop.stanford.edu/webrun
- Calculations can be performed on a new cluster at Stanford University, using the most recent GALPROP v54, older versions are also available
- The service is free and open to the community. Registration is required

Acknowledge by citing the web-page and the introduction paper <u>http://arxiv.org/abs/1008.3642</u> (submitted to Computer Physics Communications)

Configuring GALPROP via WebRun

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Image: Image			C Qr Google	
galprop.stanford.edu studies of cosmic rays and galactic diffuse gamma-ray emission				
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Search GALPROP web site	Search		Logout [avla	dim]
GALPROP version: 54			ired GALPROP v. 54 parameters and click 'Submit' at <u>the bottom of the form ↓</u>	
click to change	Common	Grids	Propagation Gas Sources Emission Abundance	es .
WebRun Help	Import configuration fro	om: you can us	se an example or retrieve your old run 🗘	n
Configure & Submit Common Parameters				
Help: Configure & Submit	Name	Value	Description	
First-time User Mode	Title	Plain diffusion mod	Descriptive title used to identify the run.	
Advanced User Mode	n_spatial_dimensions	2 \$	Specifies whether 2 or 3 spatial dimensions.	
Monitor Queue				
Download Results Energetic and Spatial Grids				
	Name	Value	Description	
	r_min	00.0	Minimum galactocentric radius (R) for 2D case, in kpc. Ignored for 3D.	- 1
	r_max	20.00	Maximum galactocentric radius (R) for 2D case, in kpc. Ignored for 3D.	- 1
	dr	1.0	Cell size in galactocentric radius (R) for 2D case, in kpc.	
	z_min	-4.0	Minimum height for 2D and 3D case, in kpc.	- 1
	z_max	+4.0	Maximum height for 2D and 3D case, in kpc.	
	dz	0.1	Cell size in z for 2D and 3D case, in kpc	
CR Propagation				
	Name	Value	Description	_
	D0_xx	2.2e28	The spatial diffusion coefficient divided by beta=v/c at rigidity D_rigid_br. The value at other rigidities is determined via the formula D=beta D_0xx (rho / D_rigid_br)^D_g, where D_g=D_g_1 for rigidity less than D_rigid_br, and D_g=D_g_2 for rigidity greater than D_rigid_br.	
	D_rigid_br	3.0e3	Rigidity for D0_xx formula, and also break point in case D_g_1 != D_g_2.	
	D_g_1	0.	Diffusion coefficient index below reference rigidity. See formula for D0_xx. Kolmogorov turbulence corresponds to a value 1/3.	
Please remember to cite GALPROP	D_g_2	0.60	Diffusion coefficient index below reference rigidity. See formula for D0_xx. Kolmogorov turbulence corresponds to a value 1/3.	î ⊥ ↓
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Interactive interface for parameter entry

Parameters are validated to avoid misconfgured runs

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